

Neutron physics: *past & future*

Ecole Joliot-Curie 2015

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Part I: History & first applications

Part II: Present day: from UCN to medicine

The neutron : *early years*

1909,1919 (and 1933!): E.Rutherford

1932: J.Chadwick, then Heisenberg

1934: F.Joliot and E.Fermi

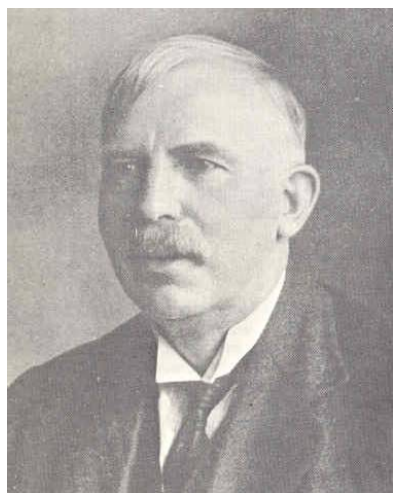
1935: I.Yukawa

1939: O.Hahn and L.Meitner

1960's: Kendall, Taylor & Gell-Mann

Why wasting time with history ?

1. « History is a lie, but a lie that isn't disputed anymore. »
Napoleon Bonaparte
2. We can learn a lot from those instructive « mistakes »
(committed by Nobel Price winners !)
3. « Old » thematics are still alive:
 - how to distinguish n/γ ?
 - J.Chadwick's brilliant idea (= *tracking* charged secondaries) is still in the heart of very very « modern » things like CMOS Active Pixel Detectors...

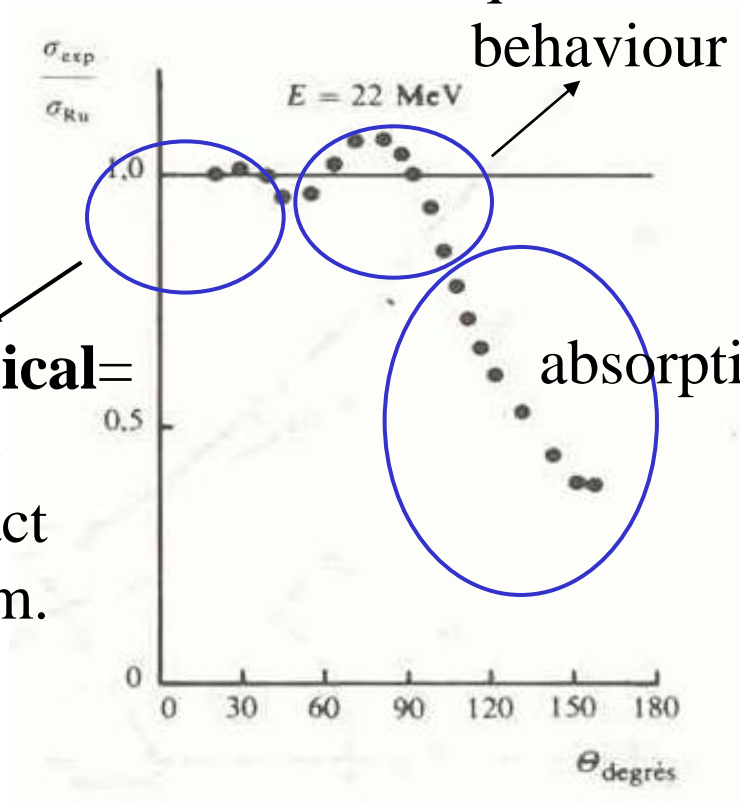


The mother of all experiments

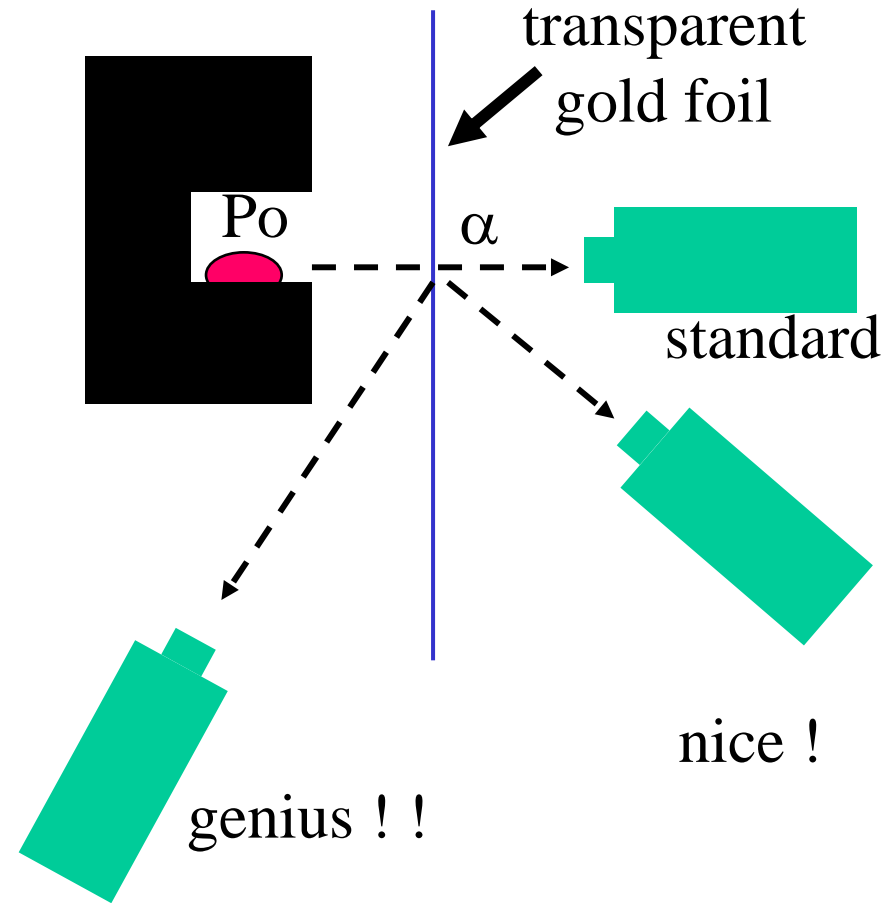
E.Rutherford, 1909

(J.J.Thomson's student...)

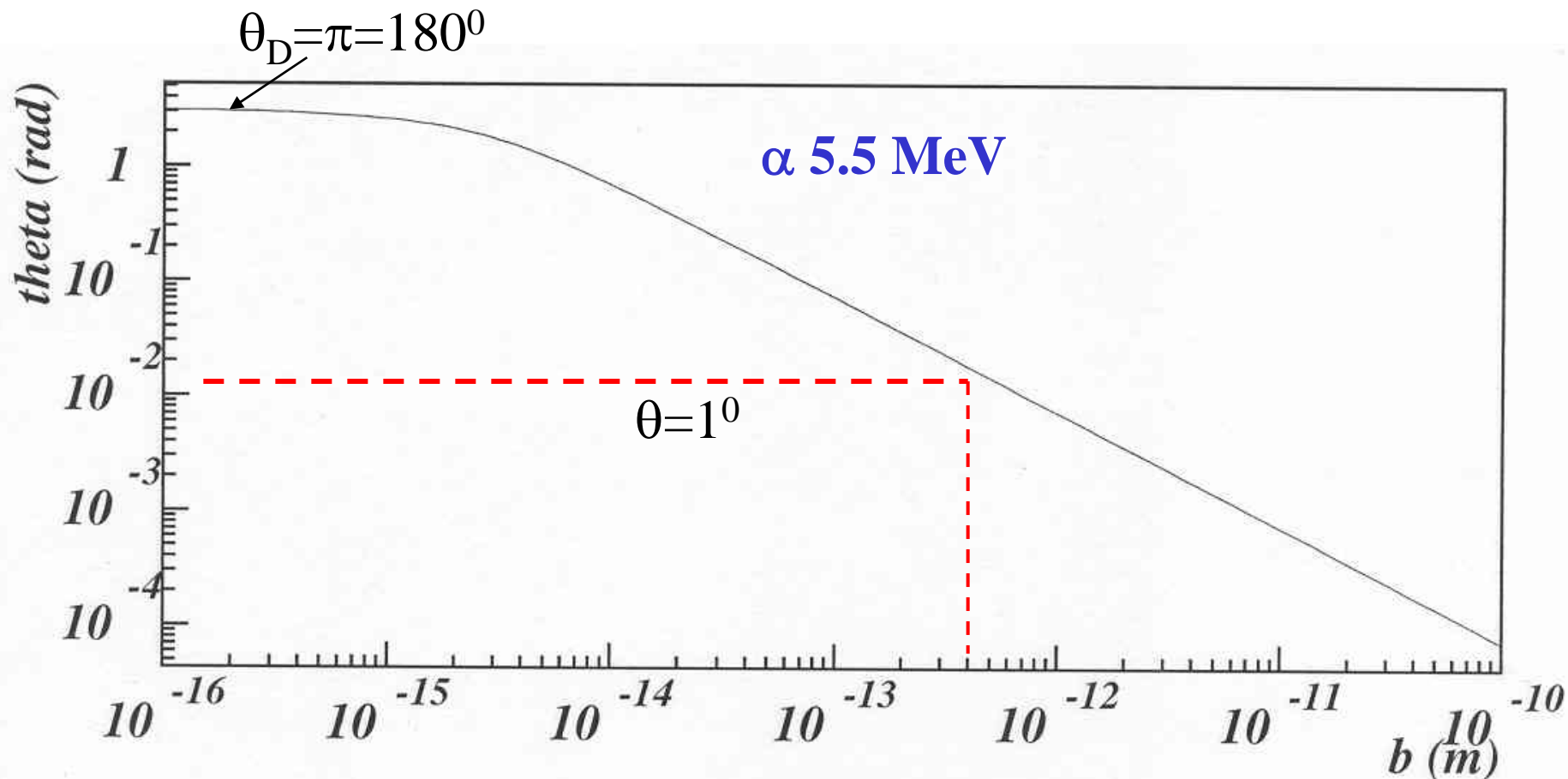
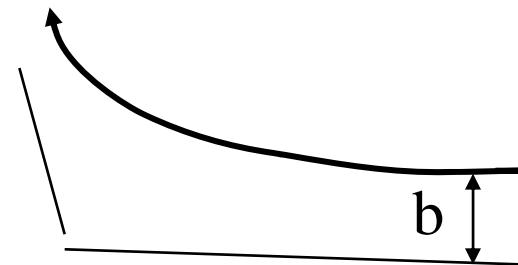
glance of quantum behaviour



classical = large impact param.



Exact calculation: $\text{tg}(\theta_D/2) = e^2/2E_\alpha b$

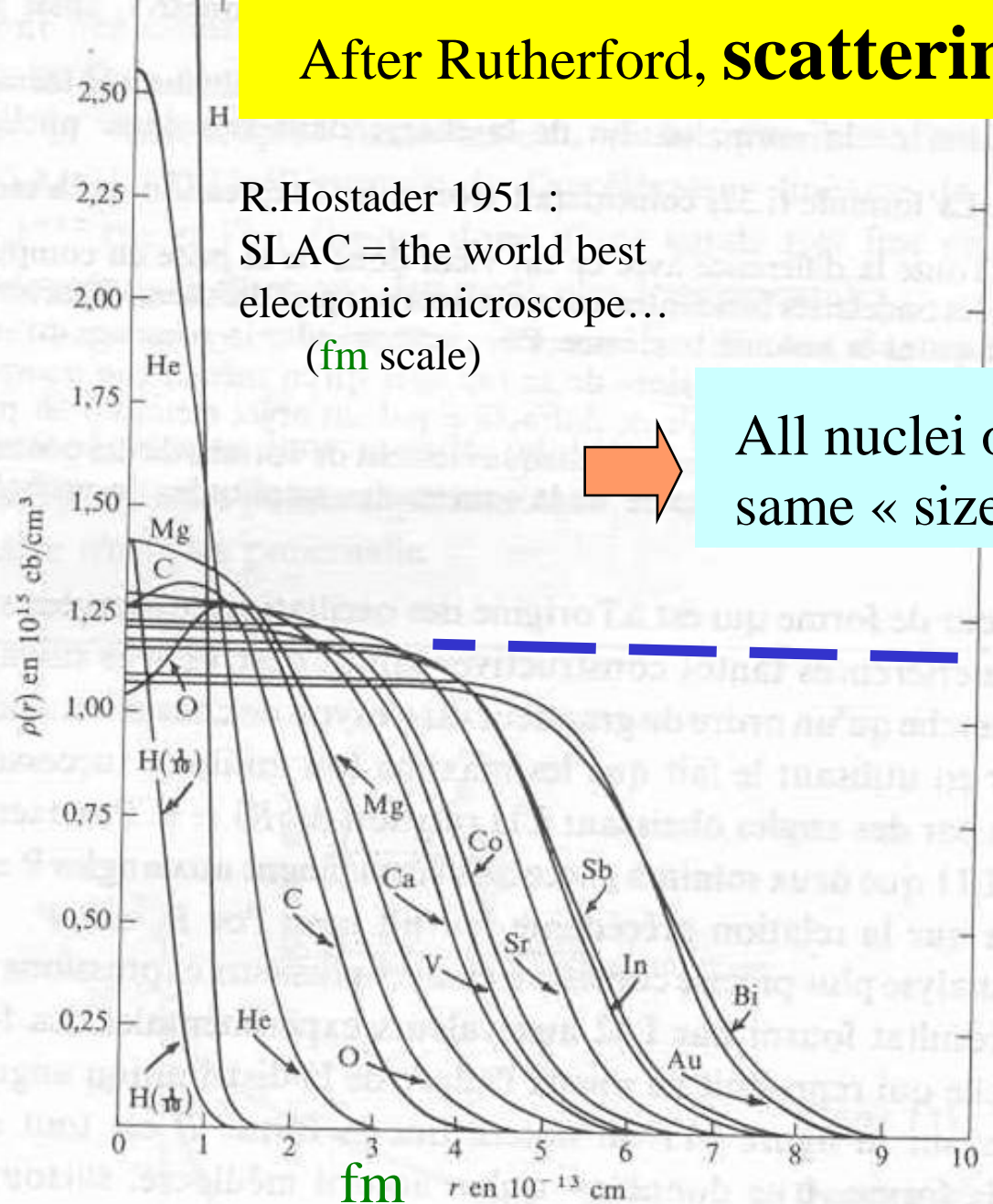


← Rare but *large* recoils →

← « matter = void » →

After Rutherford, **scattering** = the royal method

R.Hostader 1951 :
SLAC = the world best
electronic microscope...
(fm scale)



All nuclei of same « size » : $r = r_0 \cdot A^{1/3}$

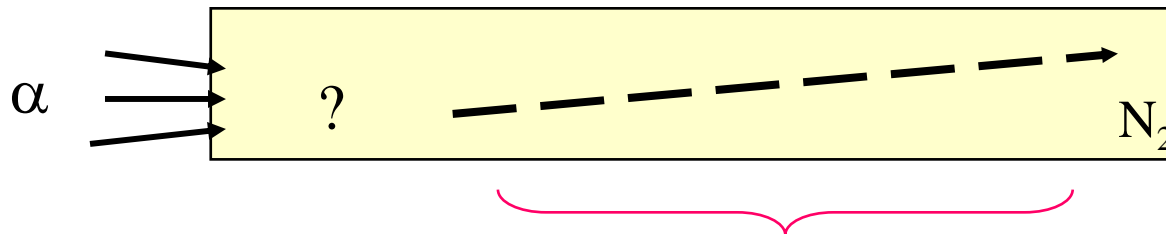
mind this nb...

$$\rho_{\text{nuc}} = A / (4/3 \cdot \pi r^3) = C^{\text{te}}$$

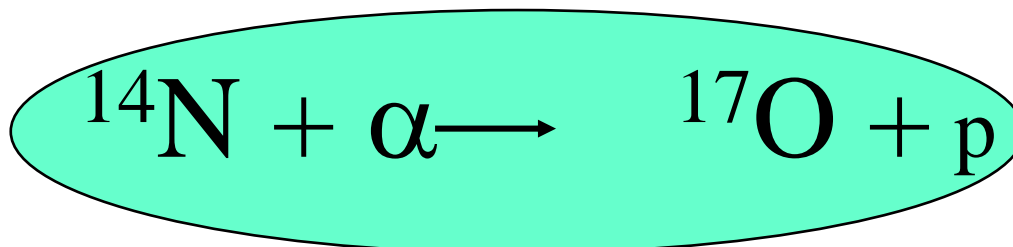
Nuclear matter has **constant density**
(=10¹⁵ times ordin.matter..)

1919: Rutherford (again!)

The very first **transmutation**:
(not Pb \rightarrow Au OK, but nice anyway...)



What *exactly* was observed ? Just a very long *range* ($\gg \alpha$) : hello *p* !!



E.R.: transmutation possible BUT very difficult (high \vec{E}) !

\rightarrow a NEUTRAL particle would be very very fine....

1920: Rutherford WELL inspired !

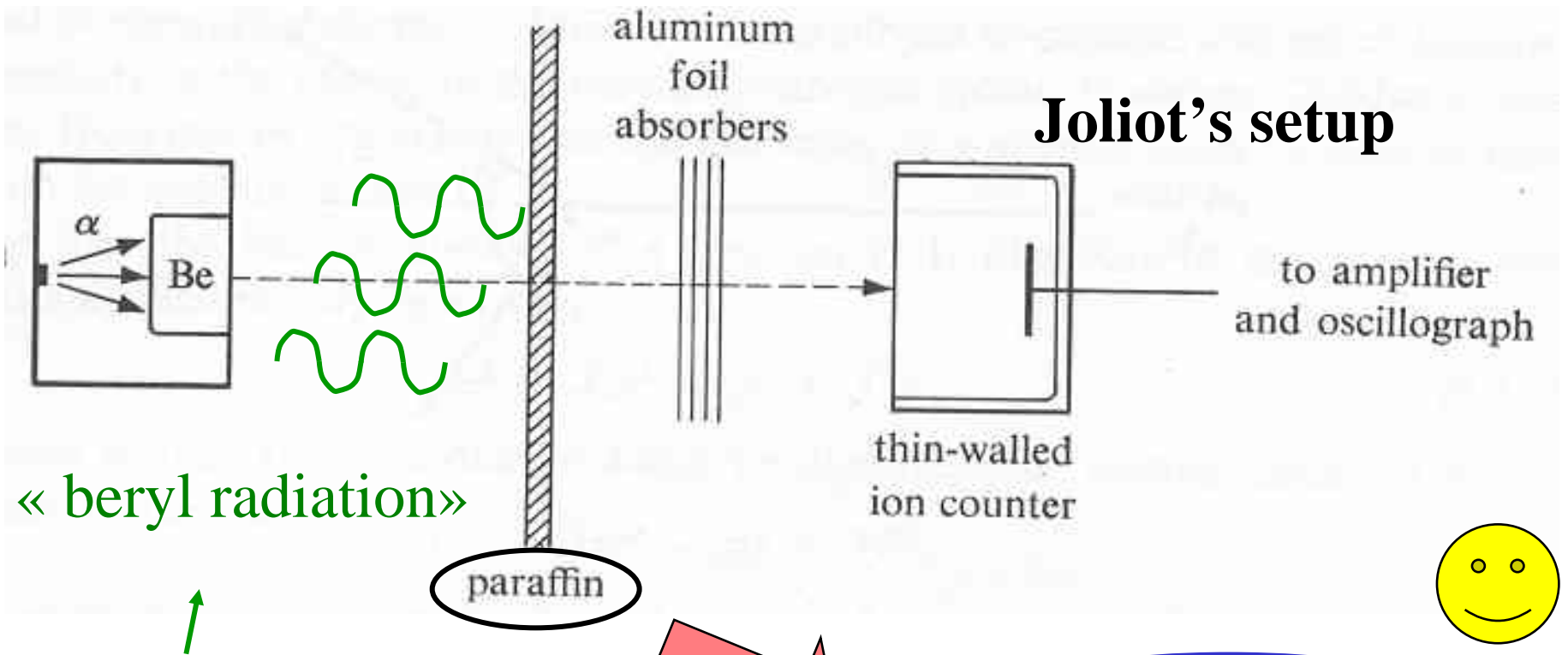
"Under some conditions, however, it may be possible for an electron to combine much more closely with the H nucleus, forming a kind of neutral doublet. Such an atom would have very novel properties. Its external field would be practically zero, except very close to the nucleus, and in consequence it should be able to move freely through matter. Its presence would probably be difficult to detect by the spectroscope, and it may be impossible to contain it in a sealed vessel. On the other hand, it should enter readily the structure of atoms, and may either unite with the nucleus or be disintegrated by its intense field.

The existence of such atoms seems almost necessary to explain the building up of the nuclei of heavy elements; for unless we suppose the production of charged particles of very high velocities it is difficult to see how any positively charged particle can reach the nucleus of a heavy atom against its intense repulsive field."

1932: Lawrence

Cited by J. Chadwick, Nobel Conf. 1935

Twelve years later: 1932

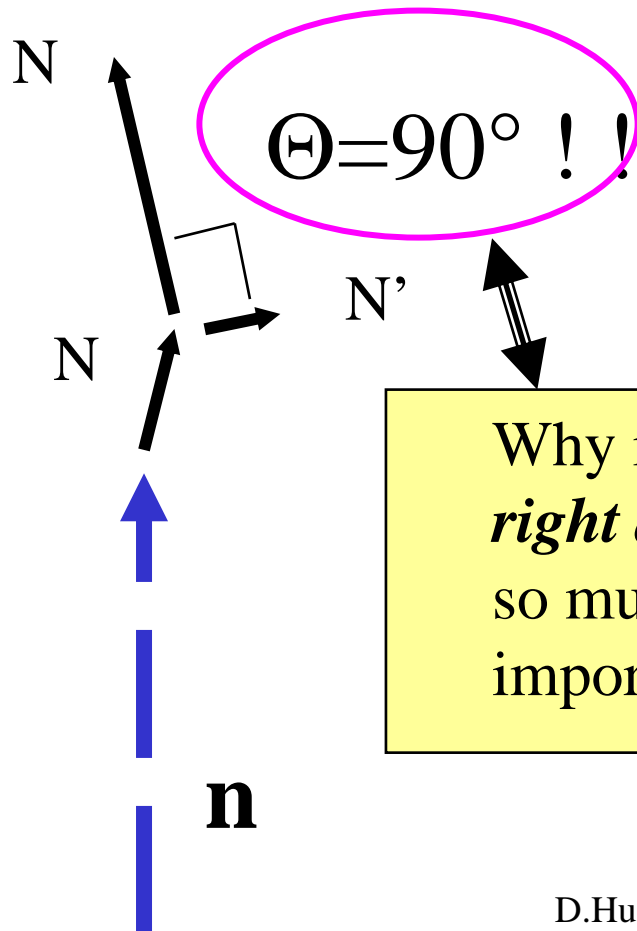


« beryl radiation »

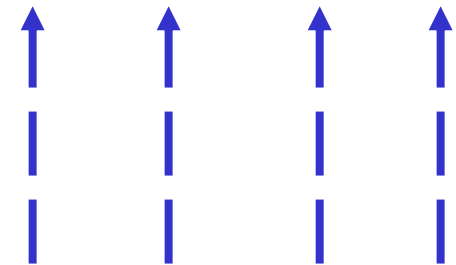
(1930:W.Bothe)

Chadwick: « let's put paraffin *inside* a Wilson chamber ! »

Chadwick's billiard game

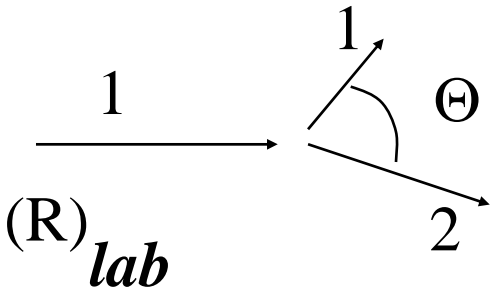
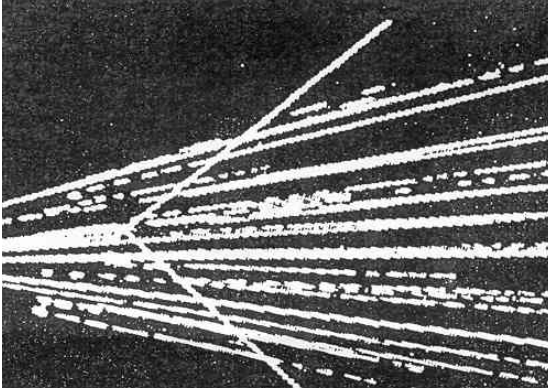


N_2 filled Wilson chamber



incident « beryll radiation »

The power of conservation laws



(1) momentum: $\vec{p}_1 + 0 = \vec{p}_1' + \vec{p}_2'$

(2) if elastic: $E_1 = E_1' + E_2'$

(3) if, in addition, $m_2 = m_1$, then:

$\Theta = 90^\circ$

this (N,N') angle is **the** proof:
 a « *beryl quantum* » is able
 to push away a *Nitrogen* ion !!

$m_1 = m_2 \Rightarrow \begin{cases} \vec{v}_1 = \vec{v}_1' + \vec{v}_2' \\ v_1^2 = v_1'^2 + v_2'^2 \end{cases}$
 $\cos \Theta = 0$
 $\Theta = \pi/2$

Chadwick's demonstration

NB: v of the postulated neutral massive particle is unknown !

⇒ Chadwick's reasoning: on maximal values



1) For hydrogen:
$$\left. \begin{aligned} m\vec{v} &= m_H \vec{v}_H + m\vec{v}' \\ m v^2 &= m_H v_H^2 + m v'^2 \end{aligned} \right\} \Rightarrow (v_H)_{\max} = v \cdot \frac{2m}{m+m_H}$$

2) Eliminate v' and set $v \cdot v_H$ to its *max* value = $v \cdot v_H$

3) Same for nitrogen ions:
$$(v_N)_{\max} = v \cdot \frac{2m}{m+m_N}$$

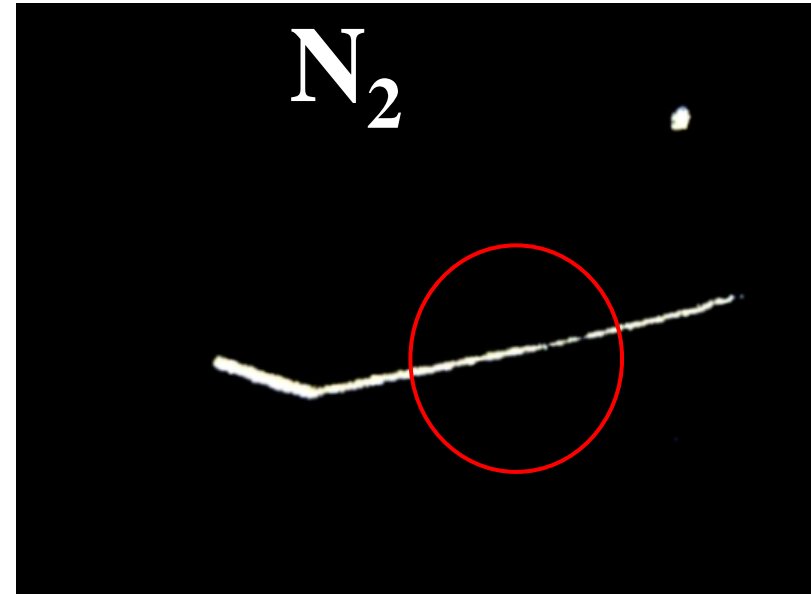
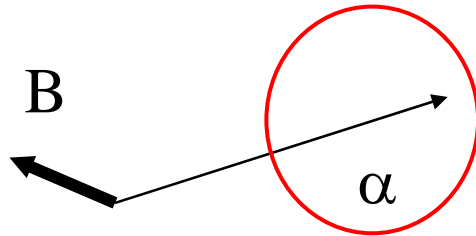
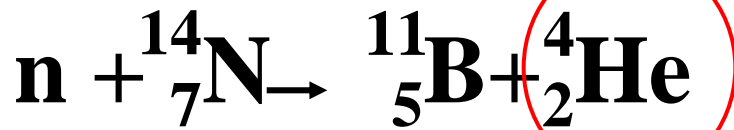
4) Ratio (eliminates v !) :

$$\frac{(v_H)_{\max}}{(v_N)_{\max}} = \frac{m+m_N}{m+m_H} \sim 7.8 \Rightarrow m \sim 0.9 m_H$$

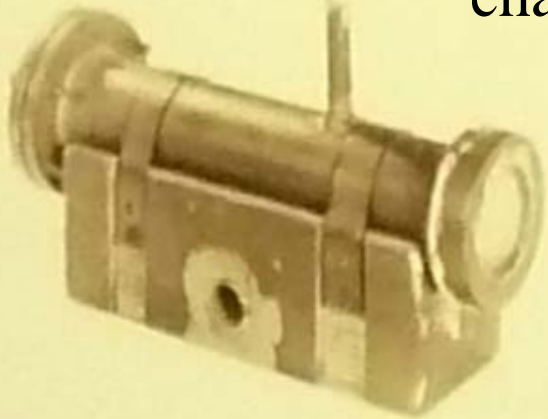
measured $37 \cdot 10^6 \text{ m/s}$
 $4.7 \cdot 10^6 \text{ m/s}$

Other nice surprises of Chadwick's billiard...

Rare event: *Transmutation* again !



This process behaves exactly as Rutherford and Fermi predicted: « much more easy » ($\sigma \times 100$) !!



chamber

What it really looked like...



paraffine disks



Top:

Chadwick's neutron chamber, 1932.

Source: Science Museum / Science & Society Picture Library

Left:

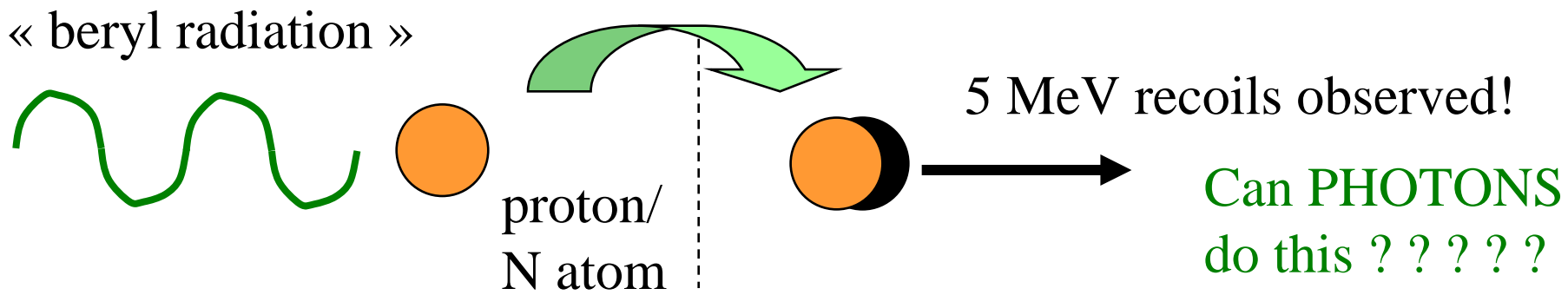
James Chadwick.

Source: Godfrey Argent Studio

1931: Joliot's historical mistake

What exactly is « beryl radiation »??

« To believe (with Joliot) that **photons** are able to eject protons out of matter is as **stupid** as trying to displace dumbbells by bombarding them with golf balls »
Ettore Majorana



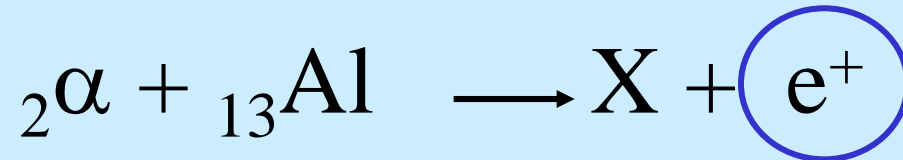
Let's calculate!

$$\begin{cases} E_p = p^2/2M \\ p = hv/c \end{cases} \rightarrow \boxed{hv = 100 \text{ MeV} !!!}$$

➡ Joliot misses the *neutron*, but...

1934

....F.Joliot does NOT miss
artificial radioactivity!



Joliot's hard work

Bombarding everything with α particles is like climbing high Coulomb barriers...but:

 Nobel Price (1935)

During 1934-1935: Fermi & al. submit everything to *neutrons* (much better probe indeed):
1) Nearly all elements undergo β^+
2) Additional surprises!

 *fission, transuranians...*
 ${}_{93}\text{Np}$ ${}_{94}\text{Pu}$...

clearsightedness
& fair-play

Cavendish Laboratory,
Cambridge.

January 29th, 1934

29. jan. of 1934

My dear Colleagues

I was delighted to see an account of your experiments in producing a radioactive body by exposure to a rays. I congratulate you both on a fine piece of work which I am sure will ultimately prove of much importance.

I am personally very much interested in your results as I have long thought that some such an effect should be observed under the right conditions. In the past I have tried a number of experiments using a sensitive electroscopes to detect such effects but without any success. We also tried the effect of protons last year on the heavy elements but with negative results.

With best wishes to you both for the further success of your investigations.

Yours sincerely

Rutherford

« a fine piece of work...which will ultimately prove of much importance »

(today: positron cameras, TEP...)

1933: Nuclear energy ?

Rutherford: « no chance ! »

- Fermi + Rutherford agree: « neutral particles should easily penetrate heavy nuclei... »
- But E.Rutherford **dismisses** the possibility of extracting the (huge) energy stored inside nuclei

1x 100 MeV ~ 0.1 nJ...

mistake is here !

SEPTEMBER 12, 1933



Lord Rutherford

Atom-Powered World Absurd, Scientists Told

Lord Rutherford Scoffs at Theory of Harnessing Energy in Laboratories

By The Associated Press

LEICESTER, England, Sept. 11.— Lord Rutherford, at whose Cambridge laboratories atoms have been bombarded and split into fragments, told an audience of scientists today that the idea of releasing tremendous power from within the atom was absurd.

He addressed the British Association for the Advancement of Science in the same hall where the late Lord Kelvin asserted twenty-six years ago that the atom was indestructible.

Describing the shattering of atoms by use of 5,000,000 volts of electricity, Lord Rutherford discounted hopes advanced by some scientists that profitable power could be thus extracted.

"The energy produced by the breaking down of the atom is a very poor kind of thing," he said. "Any one who expects a source of power from the transformation of these atoms is talking moonshine. . . . We hope in

6 jan.1939: chemist O.Hahn detects baryum in his uranyl salts...and calls Lise Meitner for help

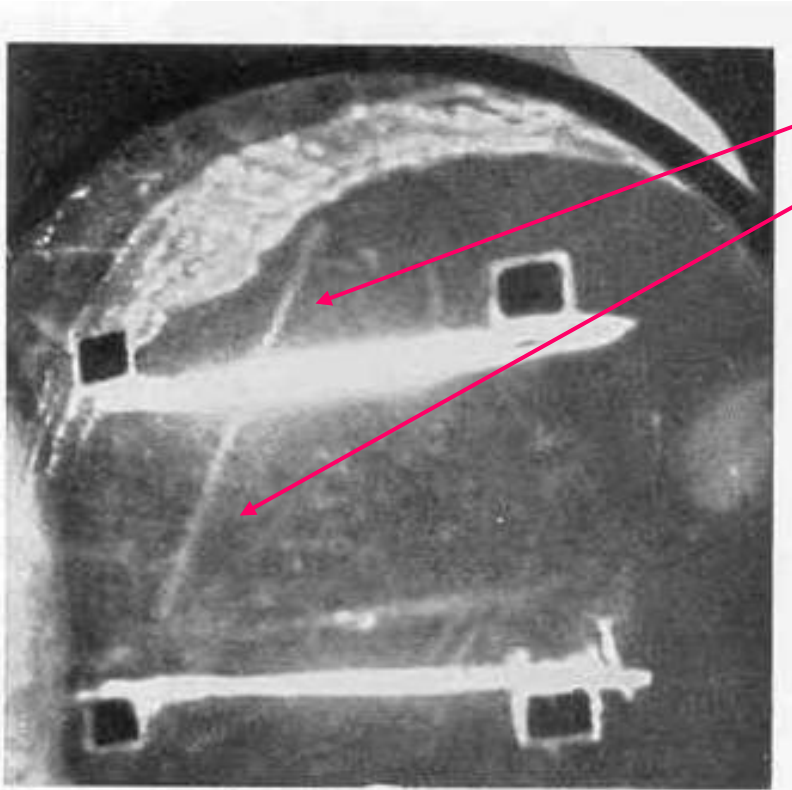


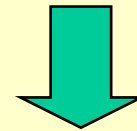
Fig. 157. Cloud-chamber photograph of HAHN and STRASSMANN's fission of a uranium nucleus upon absorption of a neutron. The two fission products leave the upper horizontal uranium plate in opposite directions with a combined energy of 160 Mev. (Photograph by CORSON and THORNTON)

FF @ 180°

What Rutherford *couldn't* know:

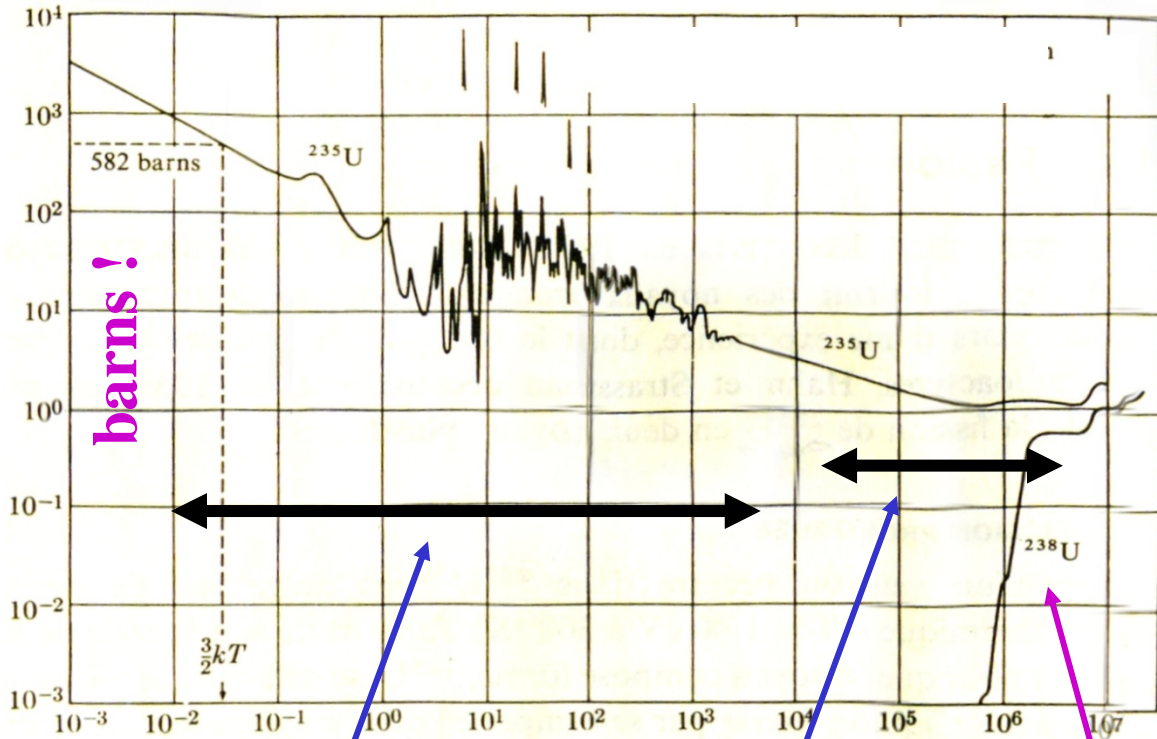
1) **Fission** happens
AND

2) $N_n = 2.416$ / fiss



THE condition for chain reaction !
(= macroscopic energy release)

Fission: all in *one* slide...



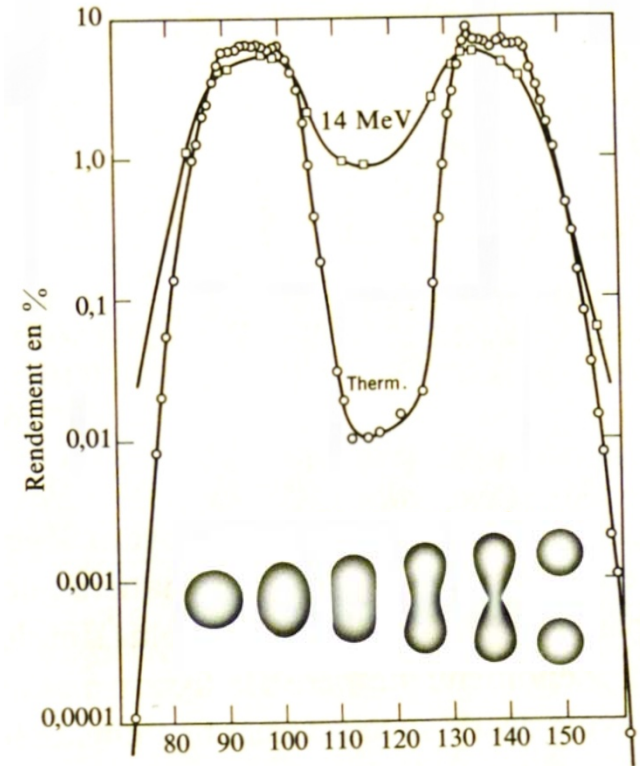
barns !

^{238}U : no!

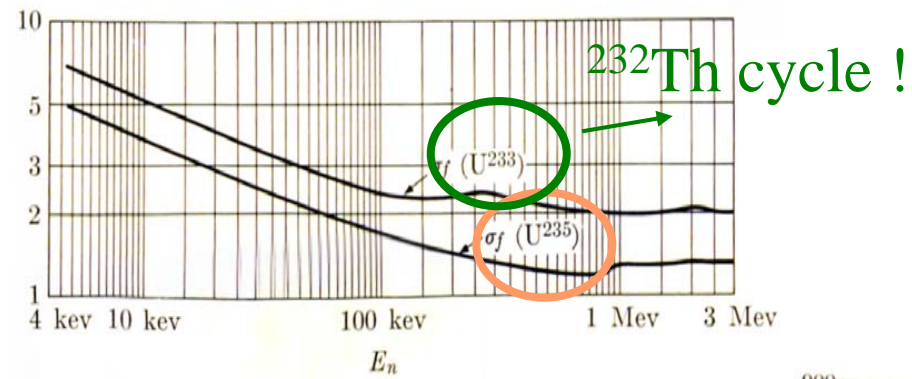
Additional n
after moderation:
 $\sigma \times 500 !$

Additional n
at creation

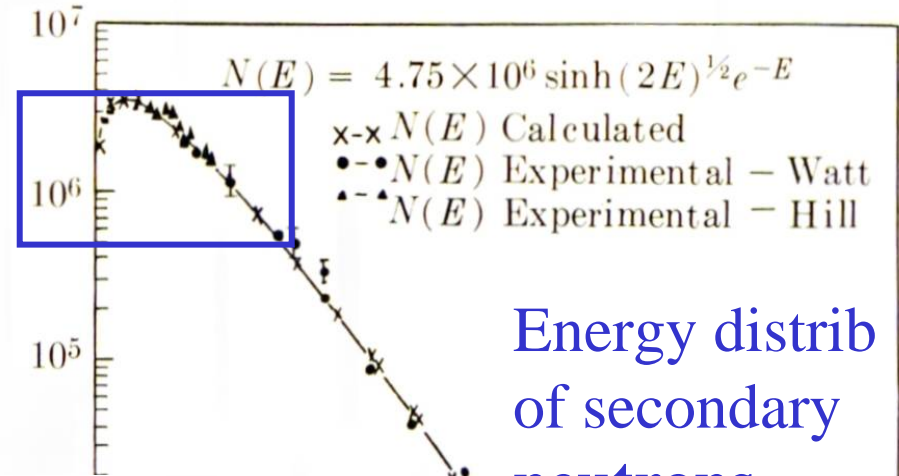
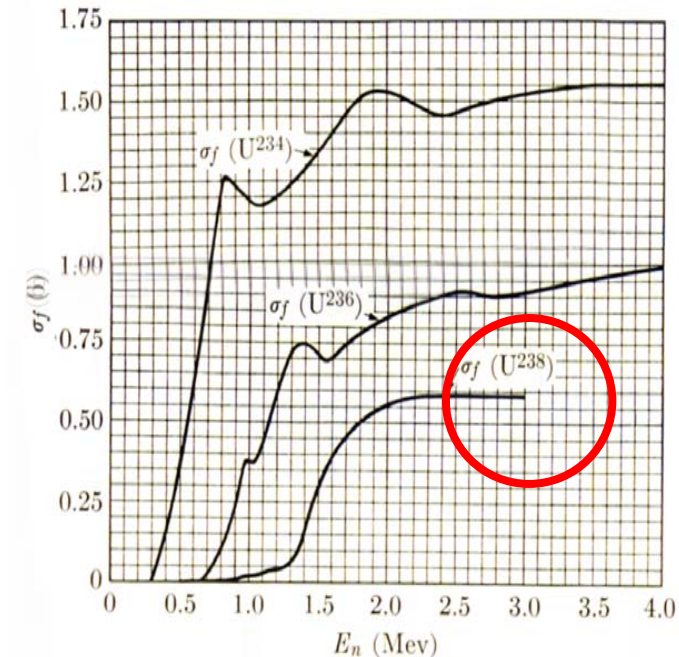
FF yield and
mass distribution



Fission: a little bit more

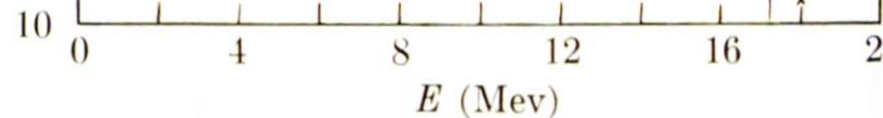
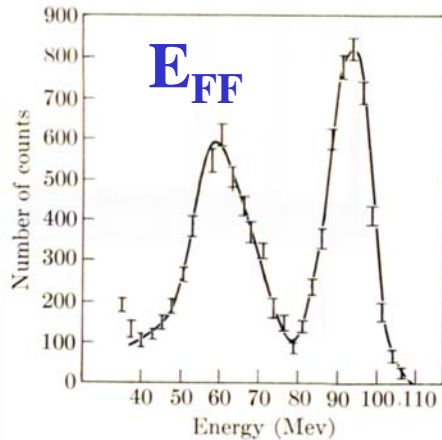


(a)



Energy distrib
of secondary
neutrons

$\langle E \rangle \sim 2 \text{ MeV}$



The neutron is a quantum object!

a) Diffusion states:

a) De Broglie wavelength:

$$\lambda = h/mv \text{ (beams)} \quad \text{or} \quad h/(2mkT)^{1/2} \text{ (« bottles »)}$$

b) Cross sections: $\sigma \sim \lambda^2 \sim \text{fm}^2$

$$\sim 10^{-26} \text{ cm}^2 = \text{barns!}$$

b) Bound states:

a) 2-particles system: the deuteron

b) N-particle systems: shells / mirror nuclei

a) the DeBroglie neutron

$$\lambda = h/p$$

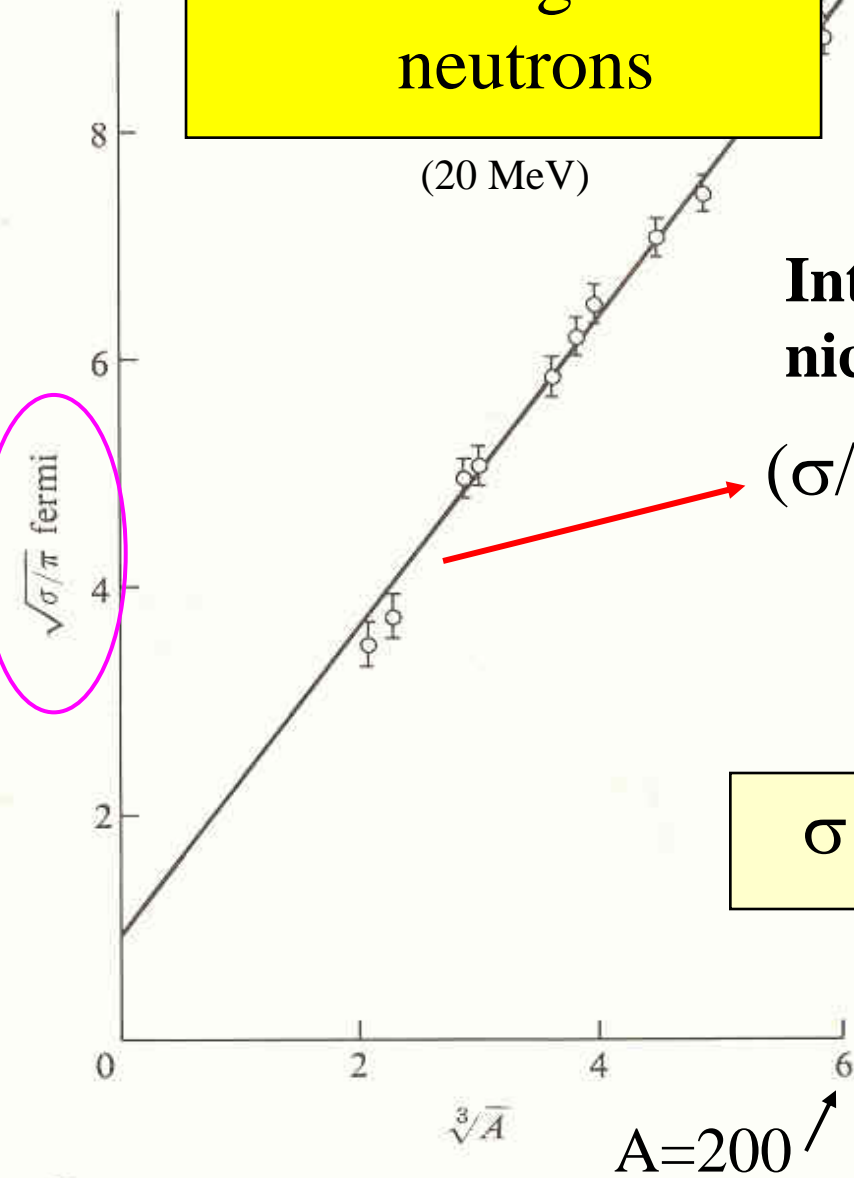
$$= h/(2mE)^{1/2} \quad (< 20 \text{ MeV})$$

$$\text{or } hc/E \quad (\text{relativistic})$$

E	v (km/s)	β_{Lo}	λ	
0.1 meV	0.14		$3 \cdot 10^{-9}$	atomic scale
1 keV	440		$9 \cdot 10^{-13}$	
20 MeV	$6 \cdot 10^4$	0.2	$6 \cdot 10^{-15}$	nuclear scale

(=6 cm/ns: limit of ToF...)

Scattering of fast neutrons



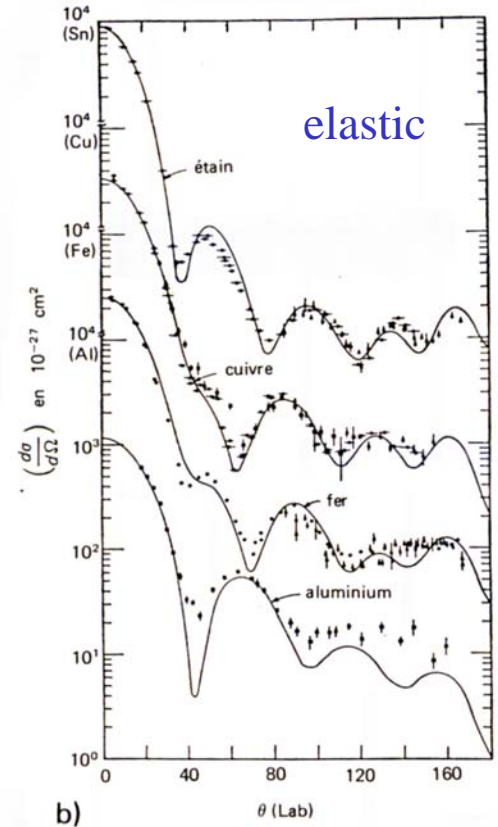
Integrated σ :
nicely linear !

$$(\sigma/\pi)^{1/2} \sim A^{1/3} !$$

$$\sim r_{\text{nuc}} ! !$$

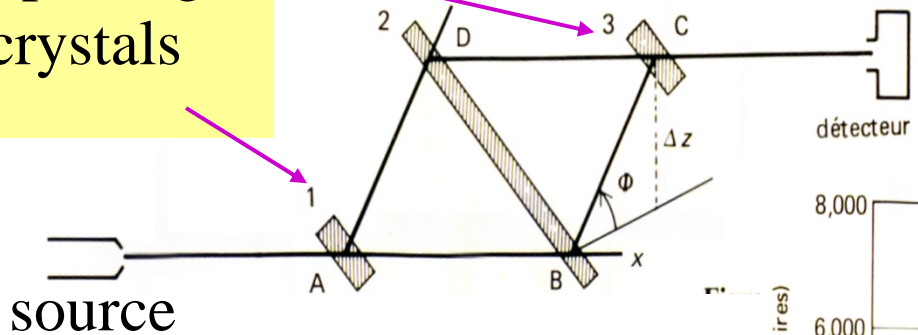
$$\sigma \sim \pi r_{\text{nuc}}^2$$

Differential $d\sigma/d\theta$:
wave mechanics!



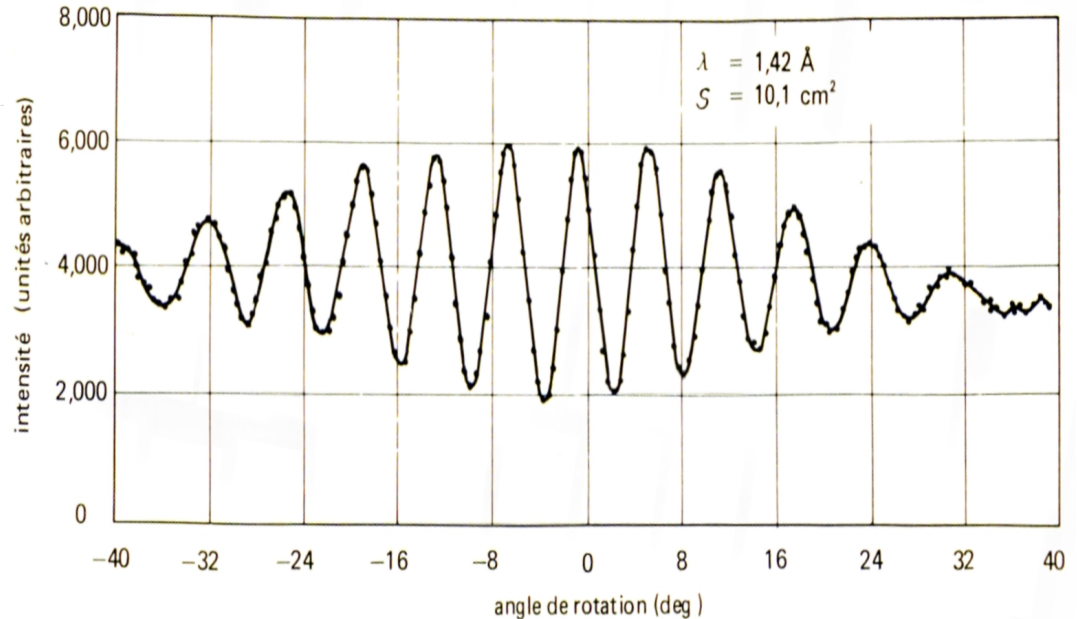
If the n is a quantic object, it should exhibit interferences...with himself

Splitting crystals



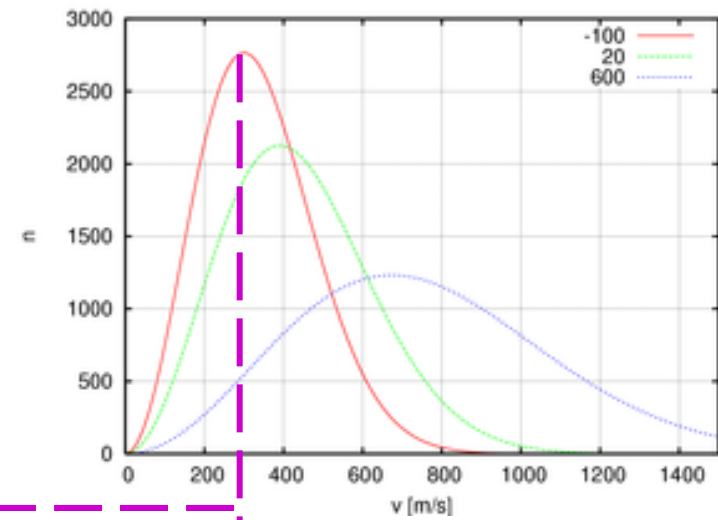
Mach-Zehnder interferom.

...and it does !



[J.L. Staudenmann, S.A. Werner, R. Colella, A.W. Overhauser, Phys. Rev. A 21, 1419 (1980)]

Reactor neutrons: thermal velocities



-Maxw-Boltzmann:

$$v_{mp} = (2kT/m)^{1/2}$$

$$= 2200 \text{ m/s} @ T=293 \text{ K}$$

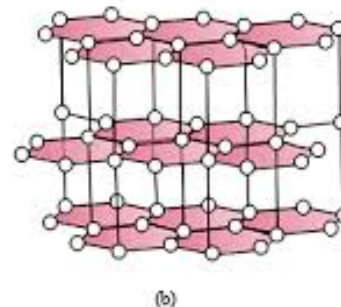
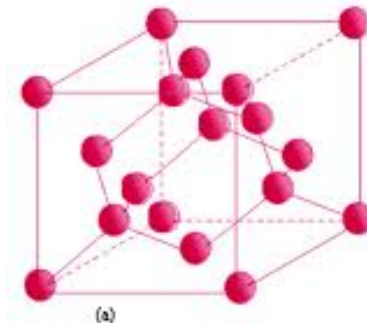
-MB+DeBroglie:

$$\lambda = h / mv_{th} \Rightarrow \lambda = h / (2mkT)^{1/2}$$

$$= 0.18 \text{ nm} @ T=293 \text{ K} \quad (\text{or } 1.8 \text{ nm} @ \underline{3 \text{ K}})$$

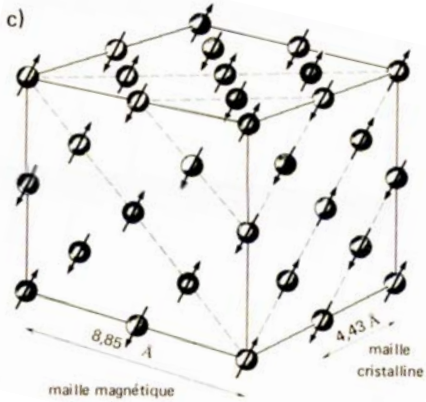


magnetic studies



↓
UCN..

Diffraction peaks \longleftrightarrow cond. matter physics



Ex: Mn^{++} has magnetic moment

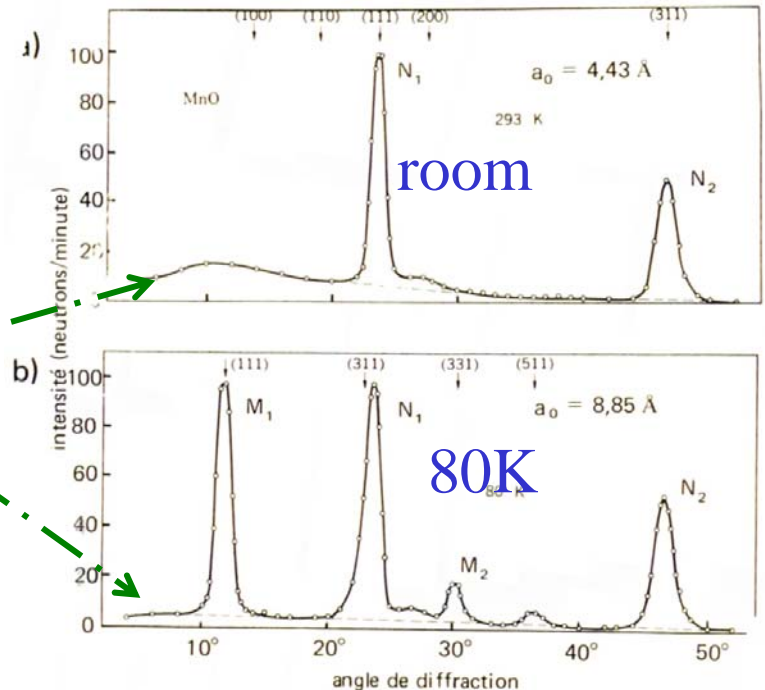
\rightarrow cubic MnO is anti-ferromagnetic

$$\vec{J}_{\text{tot}} = \vec{L}(\sim 0) + \vec{S}(\text{n}) + \vec{S}(\text{Mn})$$

-General case: J_{tot} non zero, diffusion probabilities add up incoherently.

-If polarization is prepared such as $J_{\text{tot}}=0$, two opp situations: outgoing n has $S +$ or $-$: coherent sum of *amplitudes*.

Total: sum of uncoh. backgr.+narrow peaks



Waves → interferences, diffraction peaks or...images !

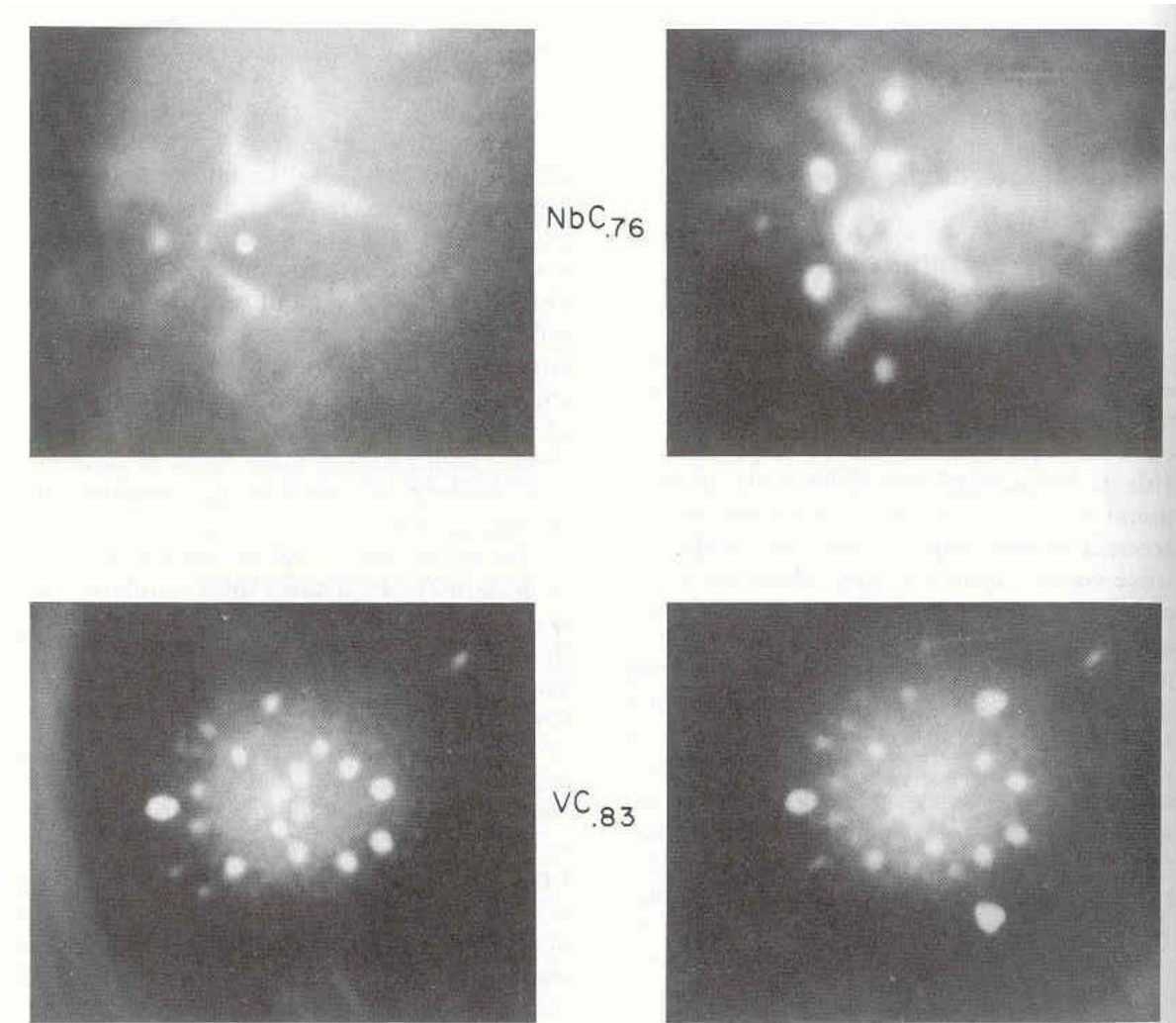


FIG. 16. The diffuse scattering pattern in $NbC_{0.76}$ due to SRO and static displacements, and the sharp Bragg reflections in $VC_{0.83}$ due to the LRO of the carbon vacancies in the crystal. (Reproduced with permission of H.

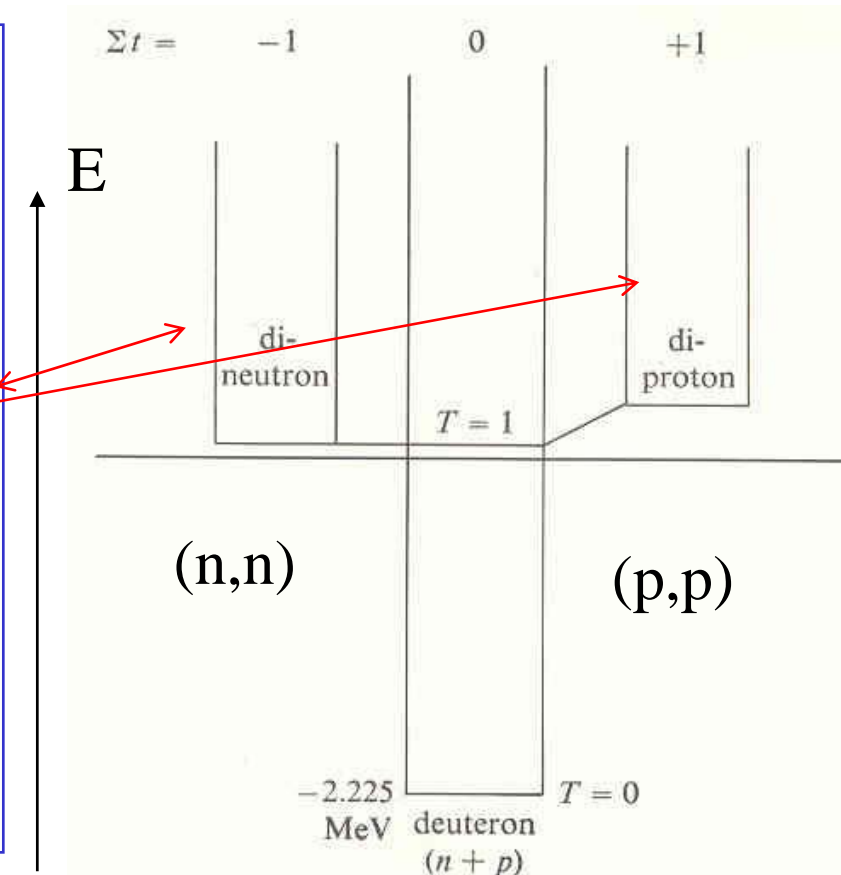
b) « royal » bound state: the deuteron

- (np) system = simplest nucleus after bare proton
- many *diffusion* states (spin dependent!) but...

ONE single bound state !!!

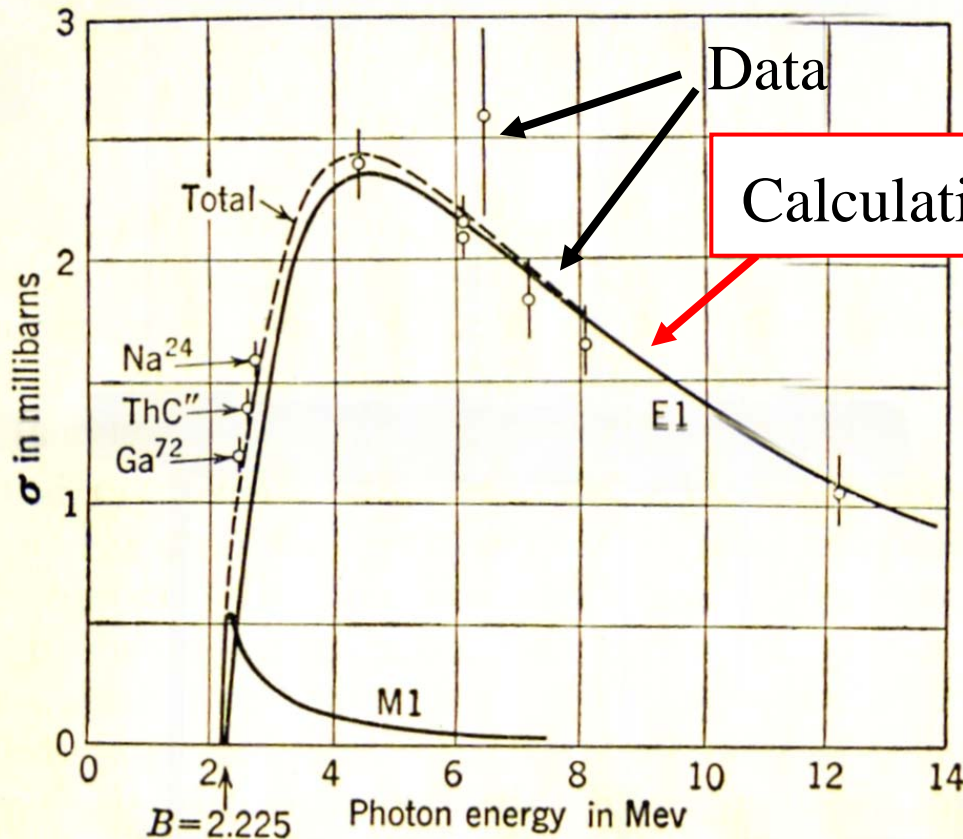
- QM prediction: Ψ must be AS
- $E_1 = -2.22$ MeV; spins: $\uparrow\uparrow$
- E_2 state = *non-binding* (spins: $\uparrow\downarrow$)
- ...but E_2 **contributes** to total σ_{diff}
- (n,n) or (p,p) states: Pauli-forbidden

Isospin symmetry :
under SU(2) Group Transf° :
p=(1,0) and n=(0,1)

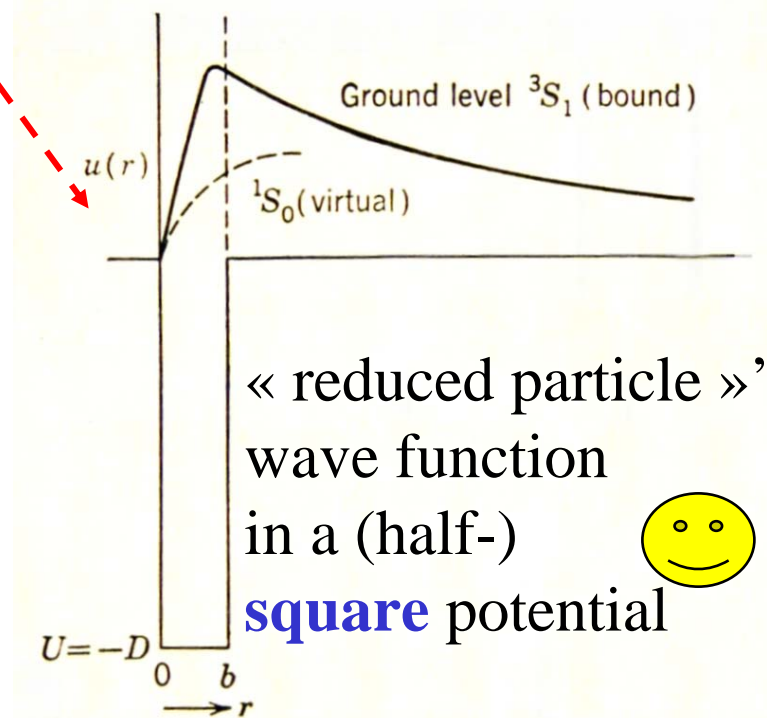


More about the deuteron (I)

Photo-dissociation: $\gamma + d \rightarrow n + p$ (inverse of fusion reaction)



↑ *Threshold = 2.22 MeV*



« reduced particle »'s
wave function
in a (half-)
square potential 😊

More about the deuteron (II)

Bound state OK ($D=-35$ MeV, $b=2.10^{-15}$ m. However! Total (n,p) at $E_n \sim 1$ keV:

$$\sigma_{\text{calc}} = 3.5 \text{ b (single bound state); } \sigma_{\text{meas}} = 20.36 \text{ b ! (big failure !)}$$

Radial pure s-wave: $\psi = u(r)/r = e^{ikx} + f(\theta)e^{ikr}/r$.

With $l=0$: $f_0 = [\exp(2i\delta_0) - 1] / 2ik = \exp(i\delta_0) \cdot \sin(\delta_0) / k$

and $\sigma = 4\pi |f_0|^2 = 4\pi \sin^2 \delta_0 / k^2$.

Lim $E_n \rightarrow 0$: δ_0 must $\rightarrow 0$, $\exp(ik\delta_0) \rightarrow 1$ and $f_0 \rightarrow \delta_0/k = -a$

$a = \ll \text{scattering length} \gg$: > 0 (bound) or < 0 (unbound) ↗

and $\sigma_0 = 4\pi a^2$ whatever > 0 or < 0 Calculation at low energy: 3.5 b !

E. Wigner: why singlet and triplet (bound st.) be of same amplitude?

With unpol. neutron beam: $\sigma_0 = 3/4(^3\sigma_0) + 1/4(^1\sigma_0)$ (as $2S+1=3$ if $S=1$)

...and obviously: $^1\sigma_0 \gg ^3\sigma_0$!

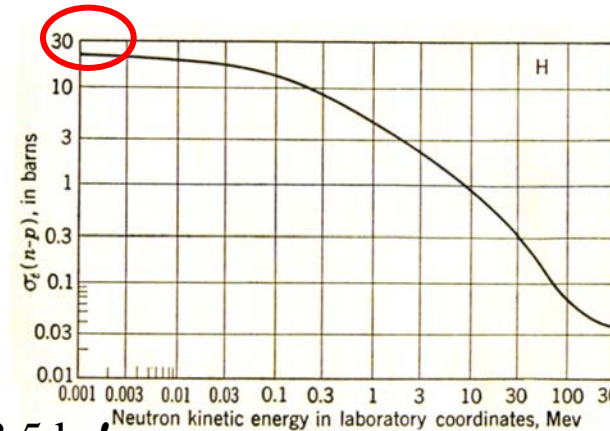
Proposed test ! orth/para- H_2 : $\sigma_{\text{ortho}} / \sigma_{\text{para}} \sim 1.5$ if $^1a = +20.10^{-15}$, but 14 if -20 fm

Exp: $\sigma_{\text{ortho}} \sim 120$, $\sigma_{\text{para}} \sim 4$ b

↑↑ ↓↑
(epitherm. neutrons).

Conclusions:

- 1) $^3(n,p) \neq ^1(n,p)$: **strongly** spin-dept
- 2) $^1a < 0$: singlet state is **unbound**
- and also: 3) $s(n) = 1/2$, not $3/2$



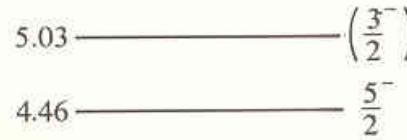
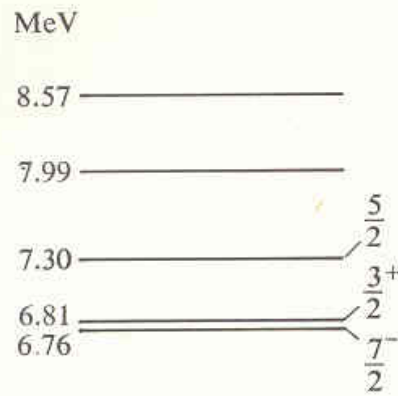
Heavier nuclei:
Heisenberg's idea

Isospin !

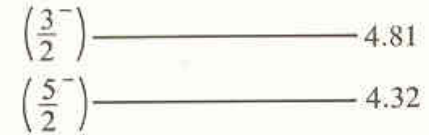
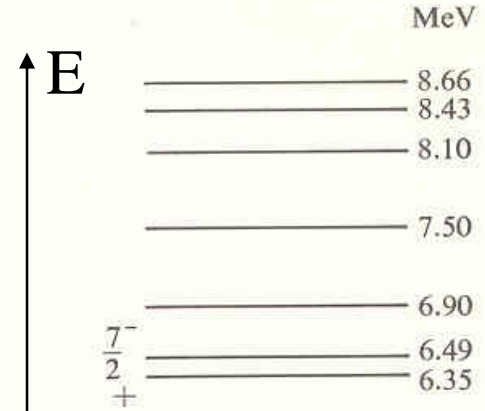
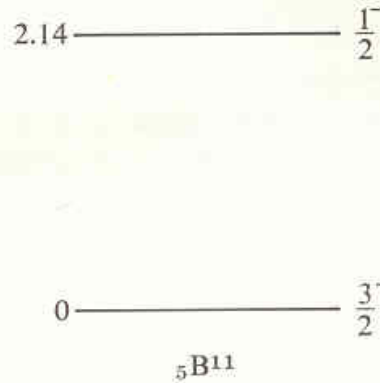


*Obvious charge
independance
of nuclear forces
(inside nuclei, just one
object: the nucleon)*

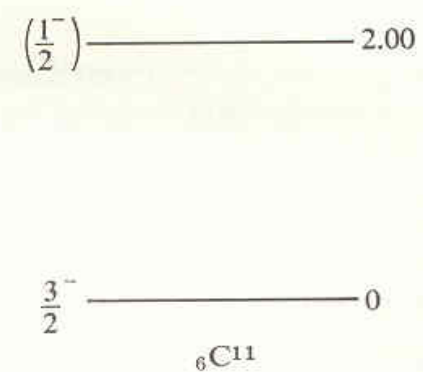
**Mirror nuclei:
same E_n and
Qu.Nbs (spin/parity)**



$^{11}\text{B} = (5p, 6n)$



$^{11}\text{C} = (6p, 5n)$



J.Chadwick, Nobel conference paper (12 dec.1935):

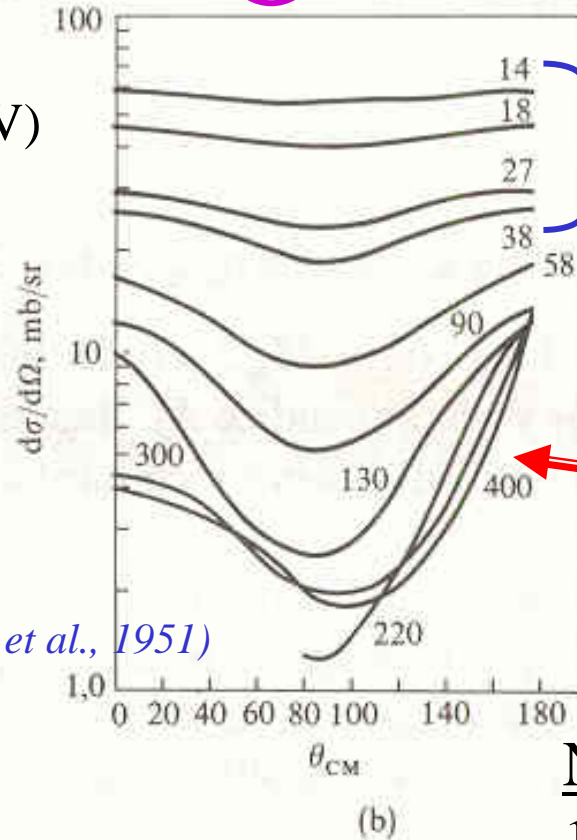
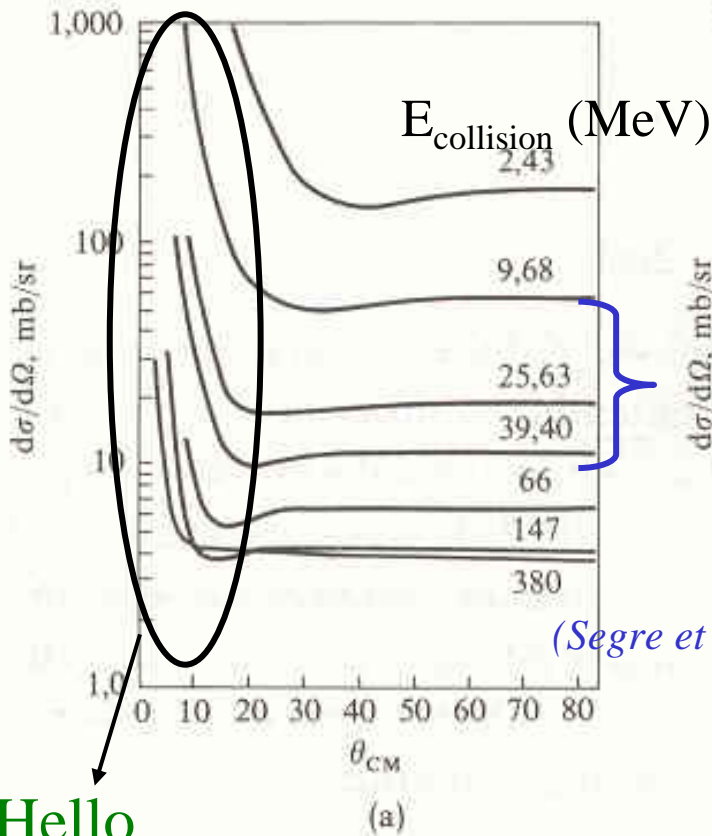
*« It appears that the interaction between proton and neutron is of the highest significance in nuc. structure, and governs the stability of the nucleus. The meagre information we have does, to some degree, support the view that the interaction is of the exchange type.
Dr.Feather and I hope to obtain more information on that by an extensive study of the collisions of neutrons and protons. »*

A step deeper: what is nuclear « force »?

→ the answer was *really* in nucleon-nucleon diffusion!

p+p

n+p



- 1) 10 mbarns/sr
- 2) $\sigma = \sigma_{\text{Ruth}}$ for p,p

3) **n ~ p** (low E)

+

4) someth.new at 130 MeV !!

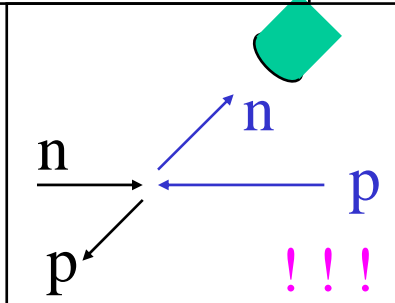
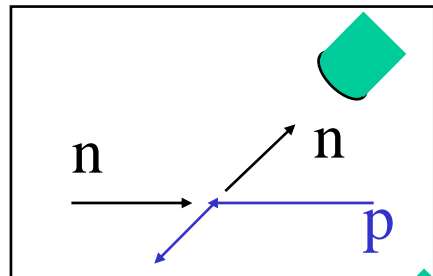
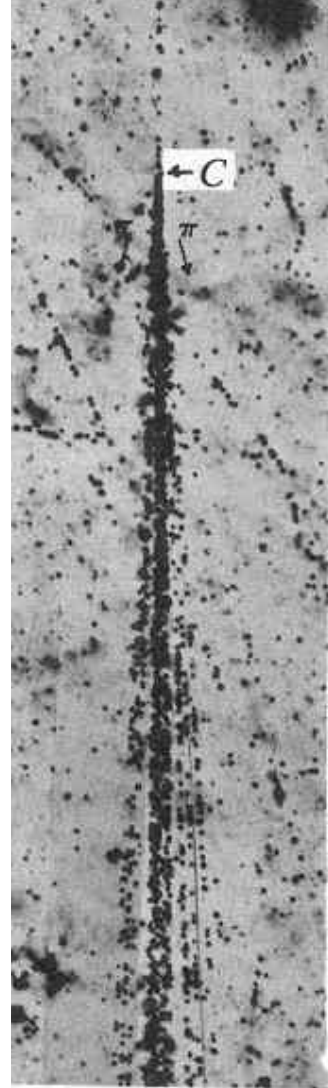
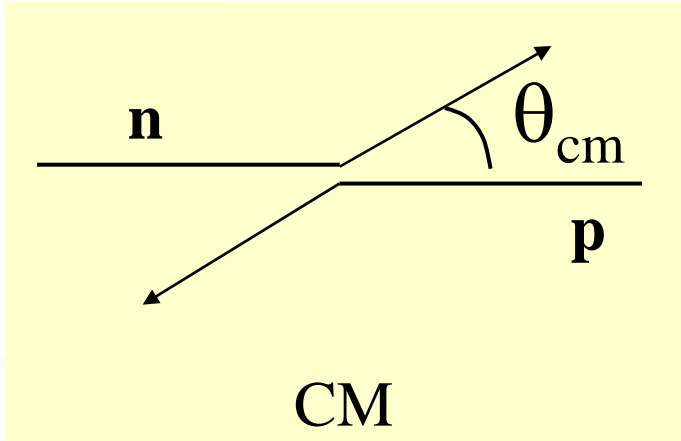
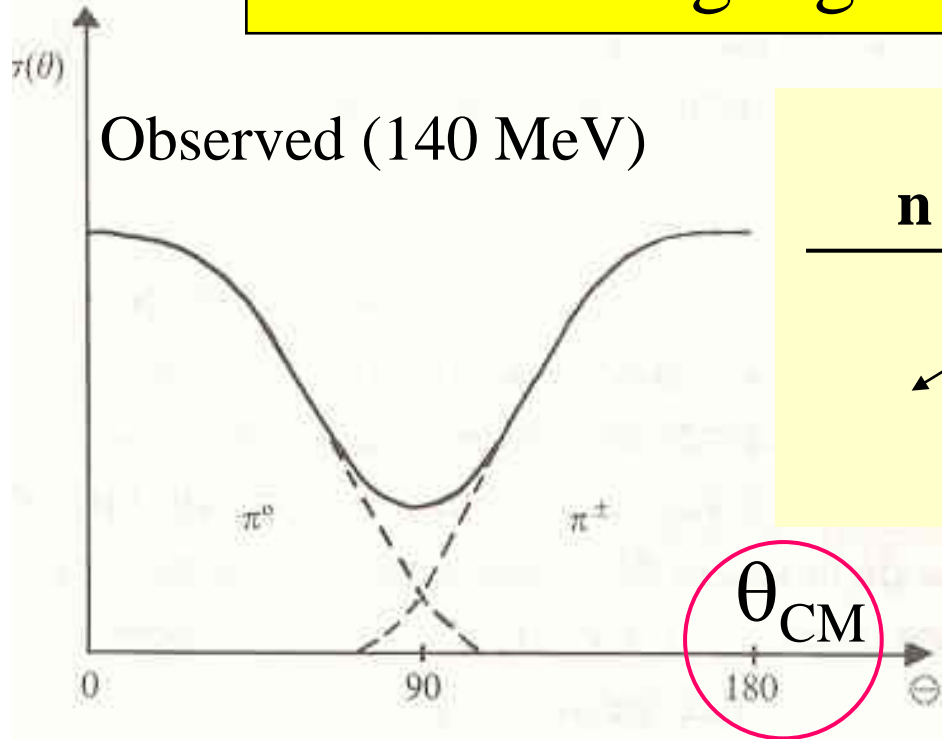
Nuc Exchange Forces !

1932: Heisenberg

1935: Yukawa

Hello
Rutherford !

Ideki Yukawa (1935) opens the road to gauge theories



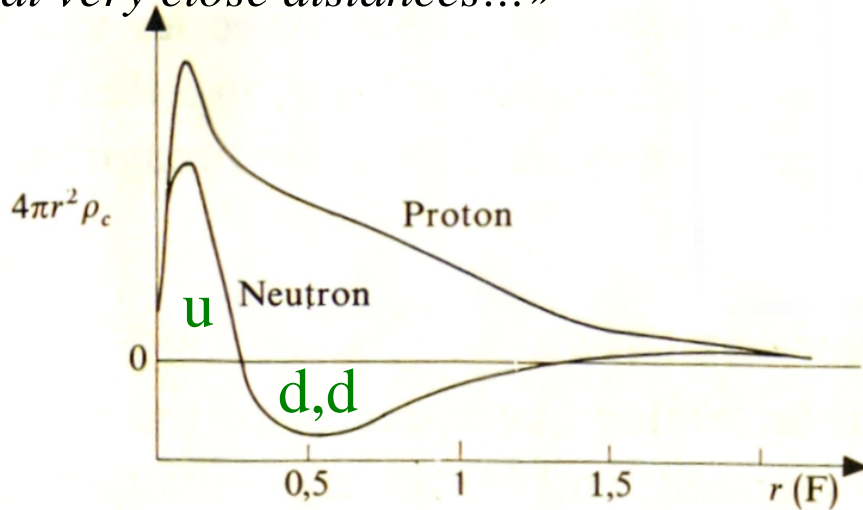
n \rightarrow p + π^- } *Boson*
 or p \rightarrow n + π^+ } *exchange*

And short range $\lambda=h/mc$:
 $\rightarrow m_\pi \sim 140$ MeV * ...

Powell * (1947)

Another step deeper: inside the neutron !

Remember Rutherford (1920) ? « *this neutral object should have **vanishing electric field** except at very close distances...* »



1961, Kendall, Taylor, Friedman
(scattering, again and again...)

Magnetic moment :

$$\vec{\mu} = \frac{1}{2} \oint \vec{r} \wedge \vec{j} d^3r$$

$\mu_p = + 2.79 \mu_N$ ($\mu_N = e \cdot \hbar / 2 \cdot m_p = 5 \cdot 10^{-27} \text{ J/T}$)
 $\mu_n = - 1.91 \mu_N$ not zero ! Net charge=0,
 but non-trivial $\rho(r)$ spatial distribution
 ➔ Quark model (1963): $p=(uud)$; $n=(udd)$

What do we do today with neutrons ?

- Fundamental Physics

1) Theory

a) Nuclear (nuc.models, halo, dynamic processes like pre-equil./evaporation, nucleon spin...)

b) Other (ν osc., UCN, quantum mechanics,...)

2) Experimental

Calibrate sources (beams) + detectors

➡ measure $\underline{\sigma}\{A(n,x)B\}$ of interest (astrophys.)

- Applied Science

a) materials (magnetism, irradiation limits,...)

b) reactors: present+future (ADR,ITER), fuel cycle

c) production of rare isotopes: ^3H , ^3He , ^{60}Co , ^{99}Tc ,..

d) biology/medical/radioprotection

Outmost important reactions for *n* detection

- ^{157}Gd : splendid gamma *signature* (ν physics)
 - ^{197}Au *activation* : precision Φ_n
 - ^{181}Hf : nice *thickness* measurement ($\sim\text{nm}$)
 - $^{108-113}\text{Cd}$: efficient *absorption* of *slow n*
 - H(n,p) elastic : historical *and* actual for fast *n*
 - $^{10}\text{B}(n,\alpha)^7\text{Li}$, $^3\text{He}(n,p)$ = fundamental reactions for :
a) slow *n detection* b) *neutron-therapy* (**BNCT**)
- } capture

..and
of course:

- ^{235}U / ^{243}Am : nuclear reactors+fuel cycle (+fiss.det)
- H/D: *moderators* => UCN, Bonner spheres...

The *n* for neutrino physics

- Discovery (1956): $\nu + p \rightarrow n + e^+$

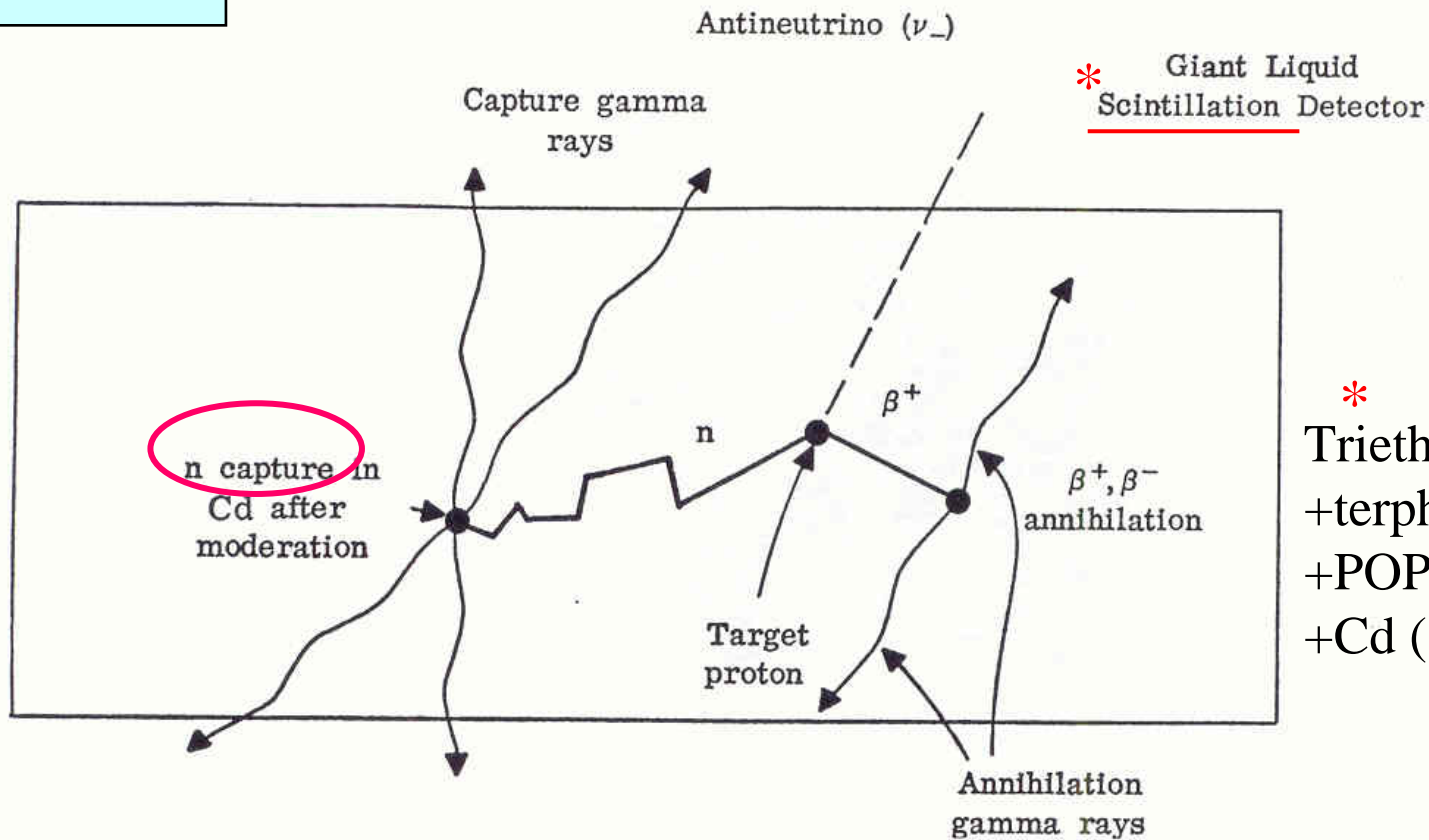
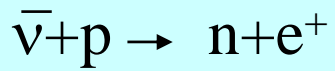
Shavannah River Reactor: detection of

BOTH *n* and positon through scintillation

- Double CHOOZ (2015): measurement of ν mass / mixing angles (oscillation setup = 2 distant det.):
= same idea (just a little bit larger...)

1956: Cowan & Reines

Savannah River



*
Triethylbenzene
+terphenyl (3g/l)
+POPOP (0.2g/l)
+Cd (1.8g/l)

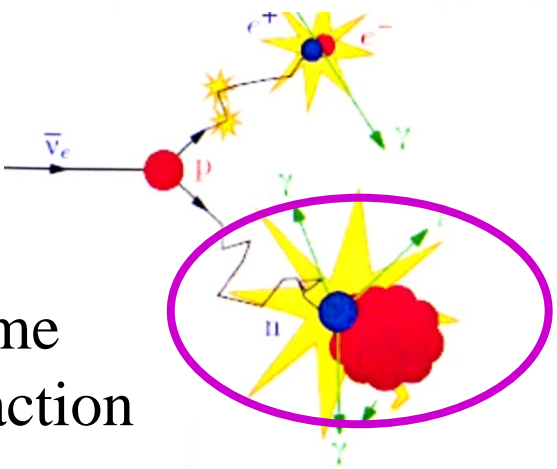
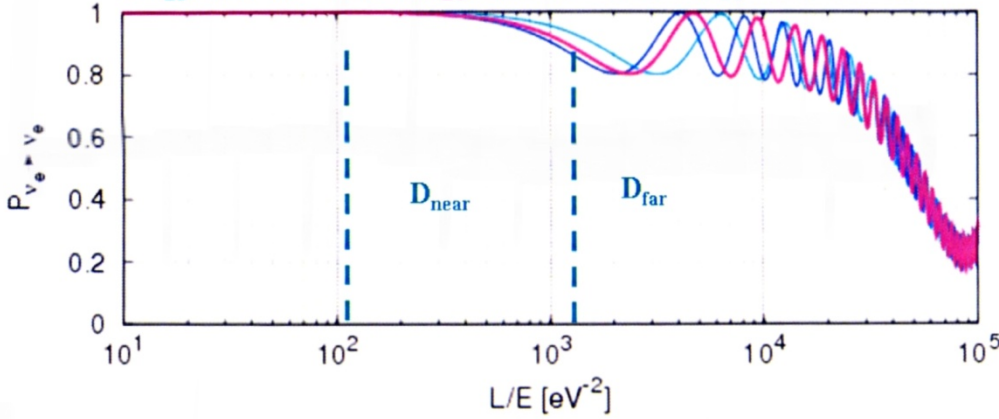
Detecting the **n** and the positron \Rightarrow proof of the 5 first (anti-) neutrinos

2015: Double CHOOZ

other context

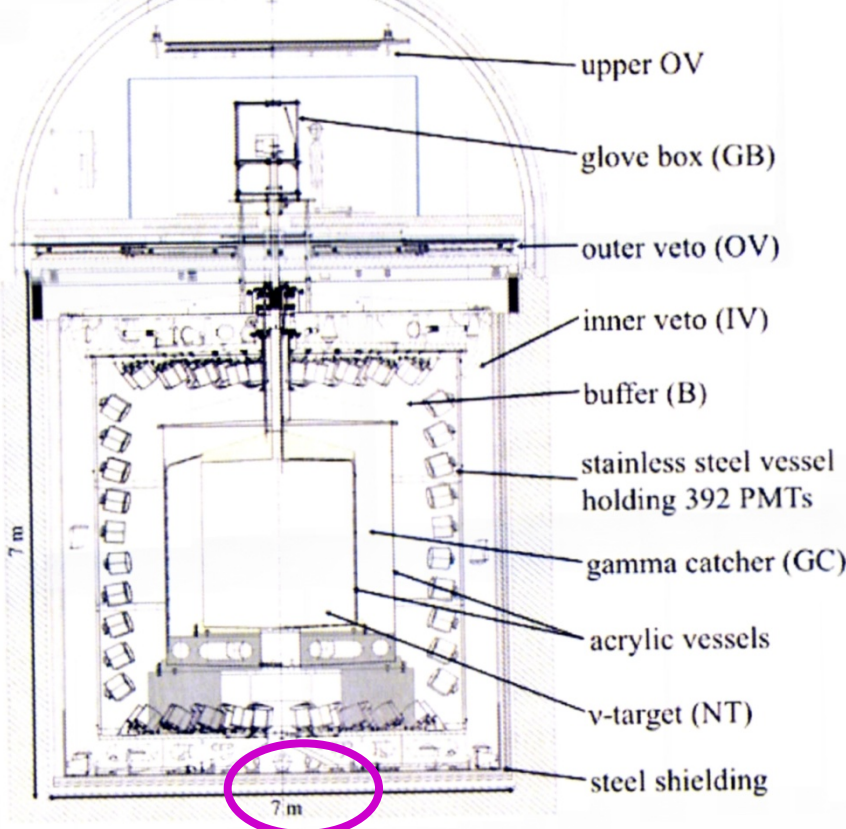
$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu1} & U_{\mu2} & U_{\mu3} \\ U_{\tau1} & U_{\tau2} & U_{\tau3} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$\Delta m_{12}^2 = 7.2 \cdot 10^{-5} \text{ eV}^2$; $\cos\theta_{12} = 0.8$; $\sin\theta_{13} = 0.23$
 $\Delta m_{23}^2 = 2.1 \cdot 10^{-3} \text{ eV}^2$; $\Delta m_{23}^2 = 2.8 \cdot 10^{-3} \text{ eV}^2$; $\Delta m_{23}^2 = 3.2 \cdot 10^{-3} \text{ eV}^2$



same reaction


bigger size..



Material activation

Very loooooooooong list of materials !

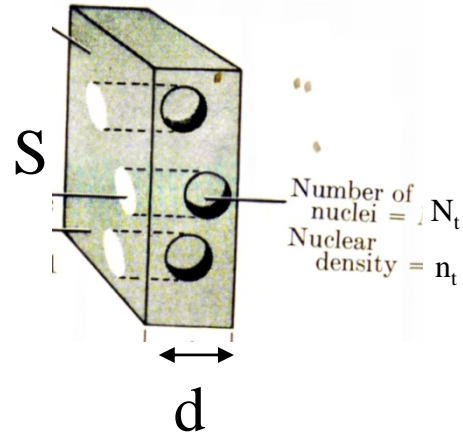
Short list of most famous targets:

Gd, Au, Al, Cu: $A(n, \gamma)B$ ( γ detectors)

Plenty of applications:

- *n* **metrology**: secondary Φ_n standards compared to ^{198}Au
- Expertise for **museums** (chemical tracks at the ppm level)
- Measuring tiny impurity concentrations (**metallurgy**)
or very small thicknesses (**microelectronics**)
-

Measuring cross sections: how?



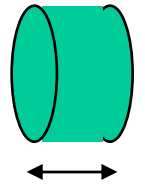
Hard spheres: $\sigma \sim \pi r_N^2 \sim A^{2/3}$

Collision proba: $P = N_{\text{coll}}/N_{\text{target}} = \sigma/S$ (thin foil approx)

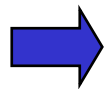
- Target density: $n_t = N_t/(S \cdot d)$ or $N_t/S = n_t \cdot d$

- Beam density $\rho(v)$ @ $v \Rightarrow$ fluence $\Phi_{\text{inc}} = \rho \cdot v$

cm^{-3} $\text{cm} \cdot \text{s}^{-1}$ $\text{cm}^{-2} \cdot \text{s}^{-1}$



$v \cdot \Delta t$



Reaction rate: $R.R. = (\rho \cdot v) \cdot (N_t \cdot \sigma/S) = \Phi_n \cdot n_t \cdot d \cdot \sigma$

From $CR_\gamma = \epsilon_\gamma \cdot R.R.$: $\sigma = CR_\gamma / (\Phi_n \cdot n_t \cdot d \cdot \epsilon_\gamma)$ (1)

Two steps procedure:

1) Choose a known material + [E1;E2] interval where $\sigma = K/v$: $RR_1 = K \cdot \rho \cdot n_t \cdot d$!!



one measures ρ (i.e. Φ) (beam calibration)

2) Take your new target, select v (or E), measure CR_2 and get the unknown σ through (1)

\rightarrow *choppers, TOF,*

The $\sigma \sim 1/v$ « law »

- Bohr (1933): for $A(n,y)B$, $\sigma(n,y) = \sigma_{\text{compound nuc}} \cdot \Gamma_y / \Gamma_{\text{tot}}$
- Breit-Wigner (1936): close to resonance (E_0, Γ):

$$\sigma(n,y) = \underbrace{(\lambda^2/4\pi)}_{\text{de Broglie}} \cdot \frac{1}{(E-E_0)^2 + (\Gamma/2)^2} \cdot (\Gamma_n \Gamma_y) \cdot (\text{add.factors})$$

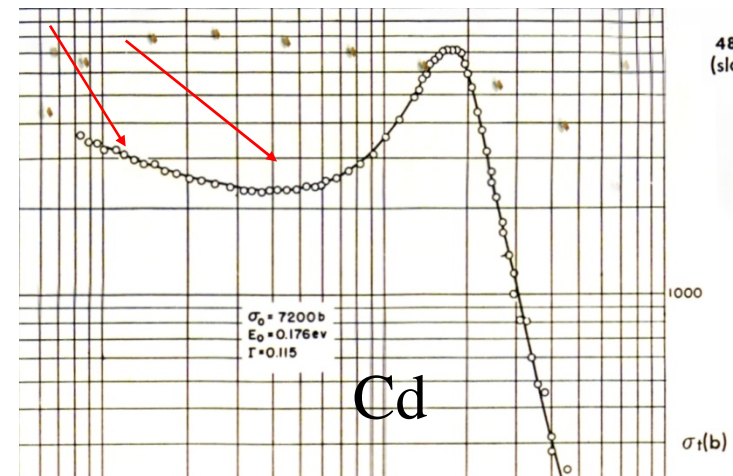
Nice feature of BW: works even *far away from resonances* !!

$$\left\{ \begin{array}{l} - \Gamma_n = \hbar / \Delta t \text{ or } \int 4\pi p^2 \cdot dp / E \sim v \\ - \lambda = h / m v \end{array} \right.$$

↓

$$\sigma \sim 1/v$$

Excellent approx. in thermal region for B, Al, Au, ...



Metrology of n beams: activated ^{198}Au to calibrate in-beam ^{10}B chambers

-Standard decay: $X \rightarrow Y$ with $N_x(t) = N_0 e^{-t/\tau}$

$$A(t) = -dN/dt = N/\tau = N_0/\tau e^{-t/\tau}$$

-Decay during creation: $X \rightarrow Y \rightarrow Z$:

$$A_\infty =$$

$$\frac{A(t)}{1 - \exp(-t/\tau)}$$

$$= N_{\text{Au}} S \cdot \sigma_0 \Phi_n$$

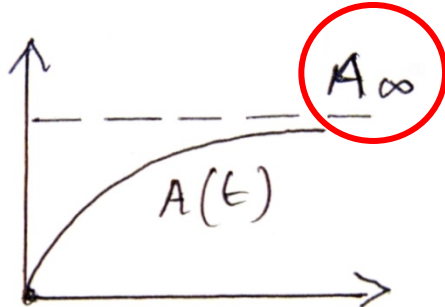
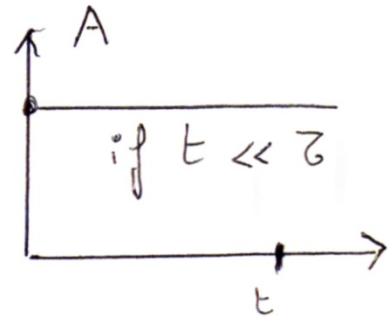
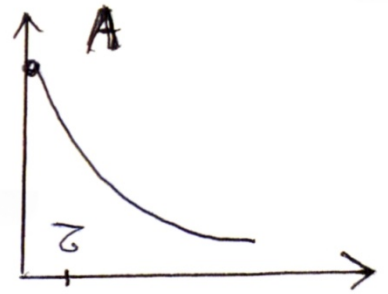
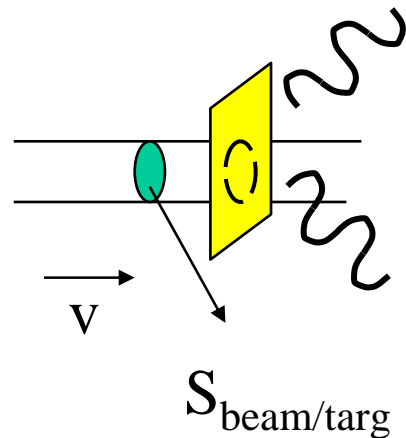
cm^{-3}

97.7 b (for **Au**)

(= $1/v$ part of σ @ $v=2200$ m/s)

$$\text{with } \Phi_n = \int v \rho \, dv = v_0 \int \rho \, dv$$

(here $\rho(v) = n$ density per unit velocity)

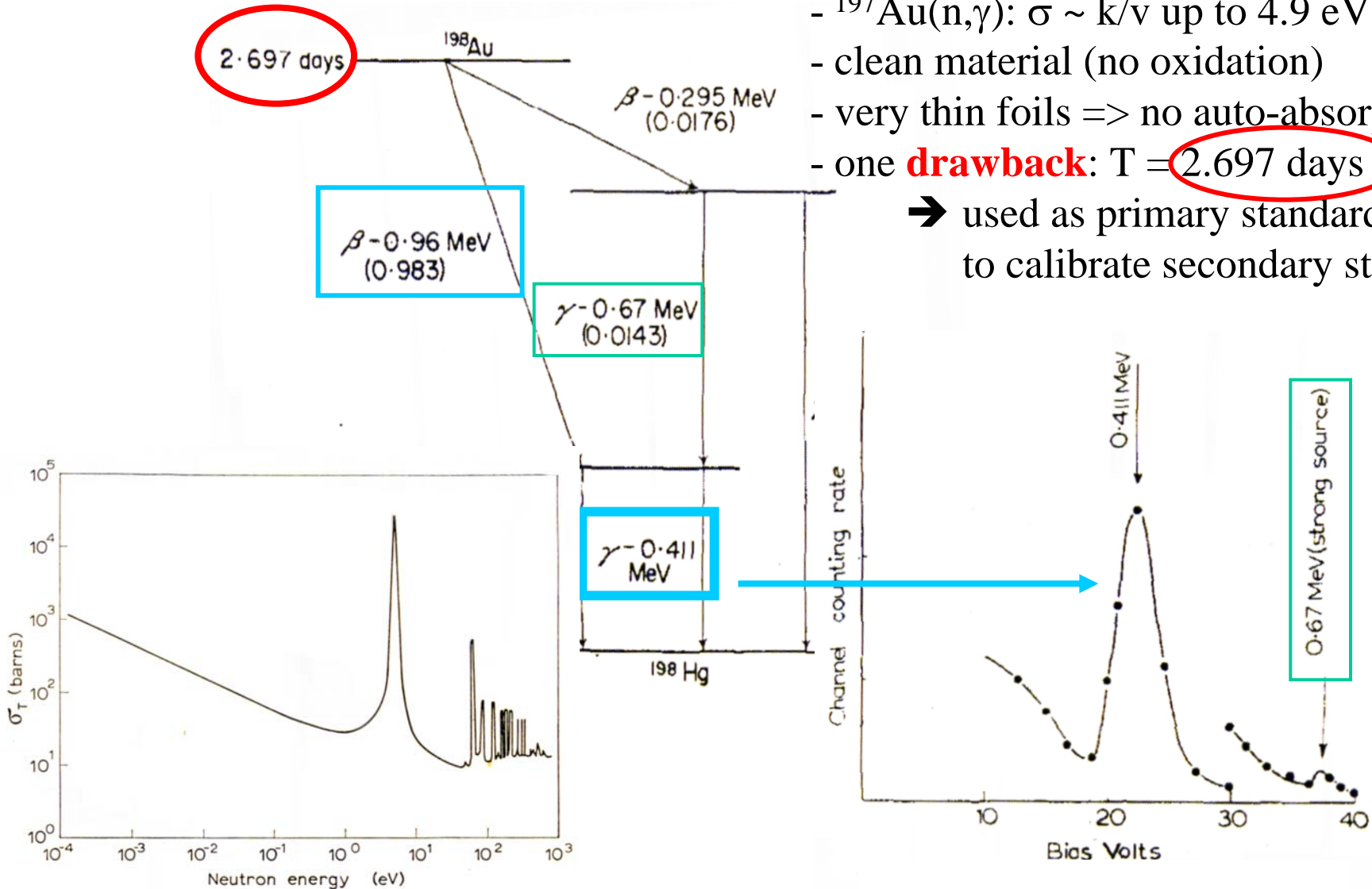


Pb: $T_{1/2}(\text{Au}) = 1/\lambda = 2.697$ days : **real time** in-beam impossible!
and $^{10}\text{BF}_3$ chambers are fast (no $T_{1/2}$!) but unaccurate.

Solution: **thin** ^{10}B layers inside prop. chambers =
good **secondary standard** (\rightarrow fast & 1% precis. after **calib!**)

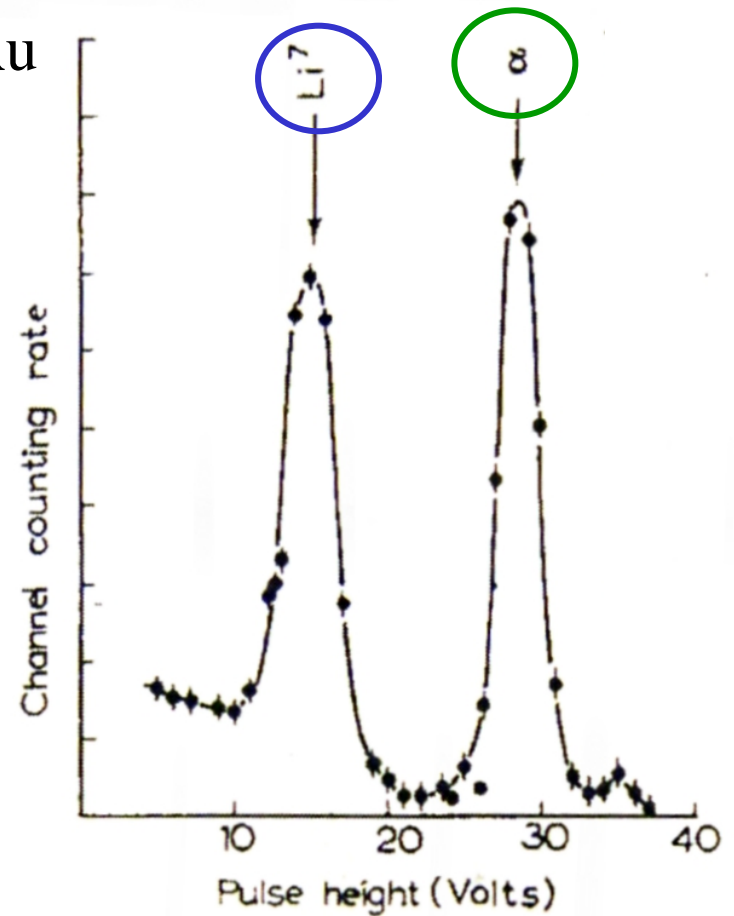
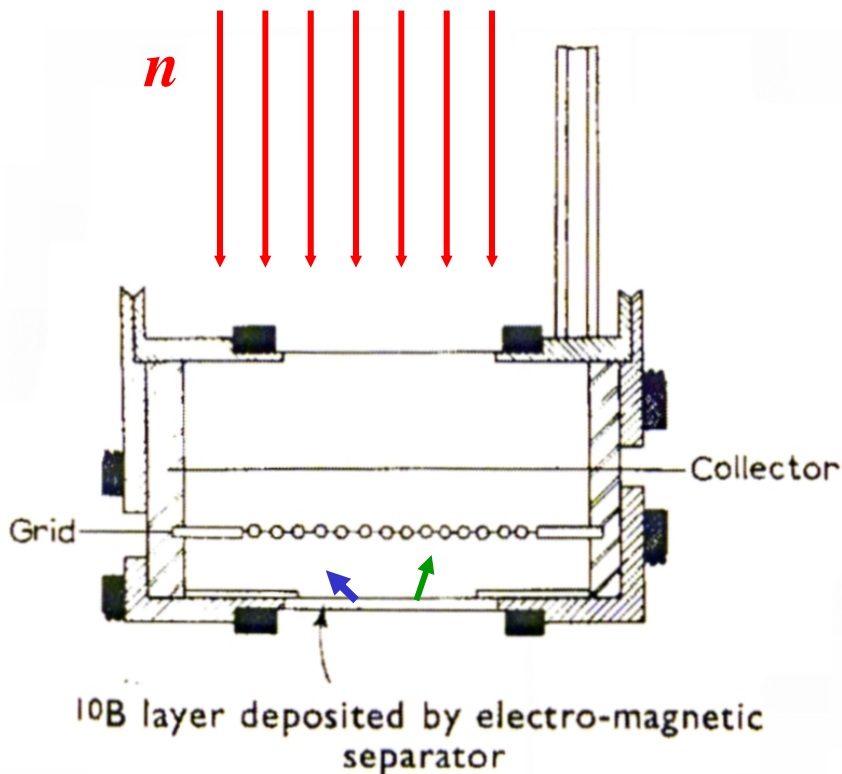
Why gold ?

- ^{198}Au : a very simple decay scheme...
 - ... plus a nice β - γ coincidence !
 - $^{197}\text{Au}(n,\gamma)$: $\sigma \sim k/v$ up to 4.9 eV reson.
 - clean material (no oxidation)
 - very thin foils => no auto-absorption
 - one **drawback**: $T = 2.697$ days
- used as primary standard to calibrate secondary standards



NOW the ^{10}B thin foil chamber is useful as a secondary standard:

- a) accurate because *thin* foil (absorpt \sim 0)
- b) fast (no decay time) => *real time* !
- c) peak height calibrated against ^{198}Au



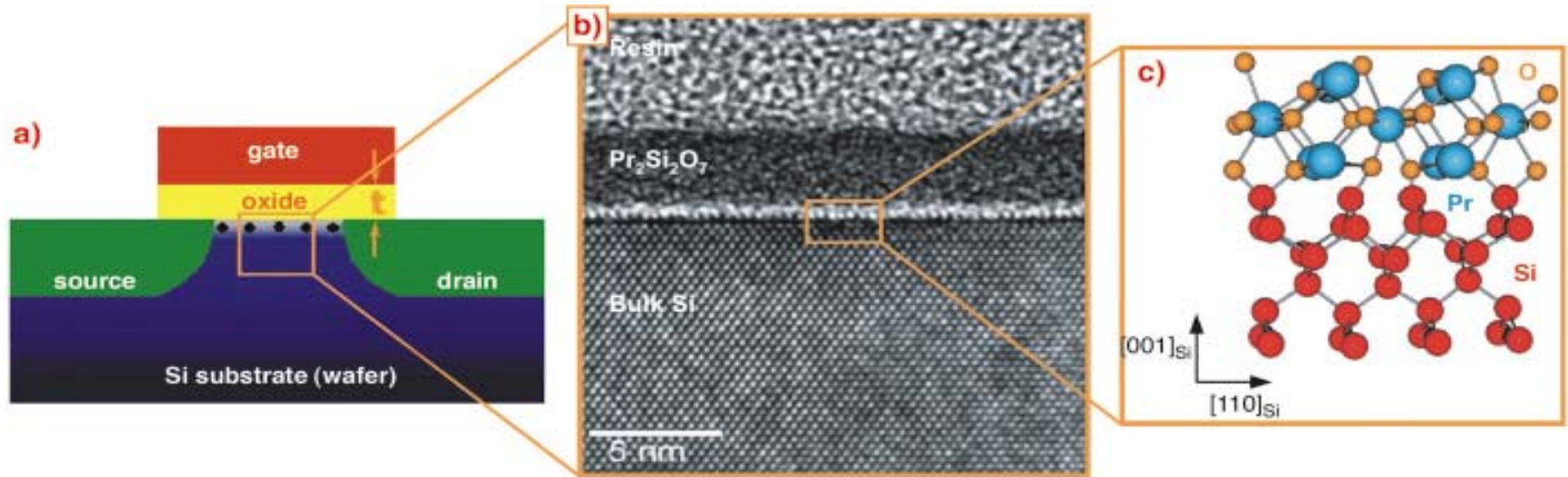
Analysis of historical paintings

- Method: Neutron activation autoradiography (NAAR)
- Complementary to XRR and macro-X-fluoresc.
- Only $4/10^{12}$ atoms sensitive to n
- Able to reveal special pigments, e.g. Mn,Cu (azurite) , As, Co, Pb (lead white), Hg(vermilion) and mainly P under opaque layers



M.Alfeld et al., Appl.Phys A (2015)179

Nanoscale transistors need « high-k » dielectrics !



Roadmap for microelectronics =

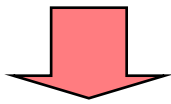
desperate run to *smallest sizes* + same performances!

→ But : just a little problem to keep $\epsilon_r \epsilon_0 S/d$ with $S \searrow \dots$

→ exotic oxides/metals (Pr, Hf) to get $\epsilon \sim \text{SiO}_2 \times 20, \times 50 \dots$
for dielectrics of 5-6 atomic layers !

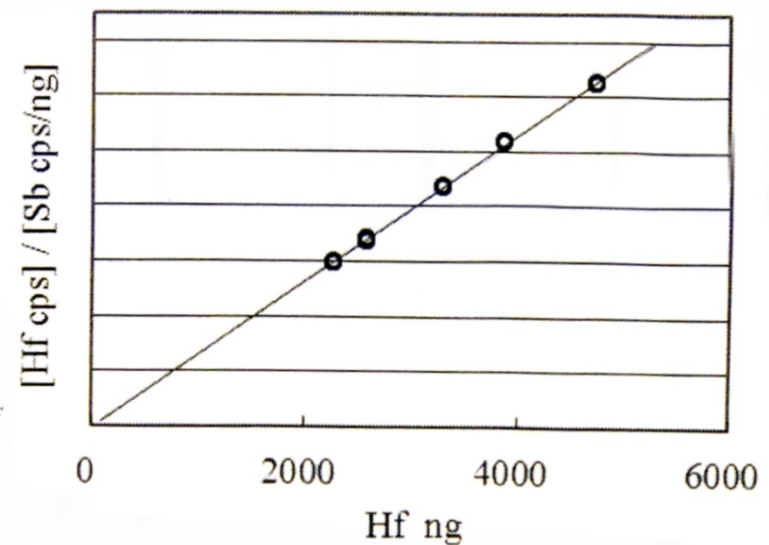
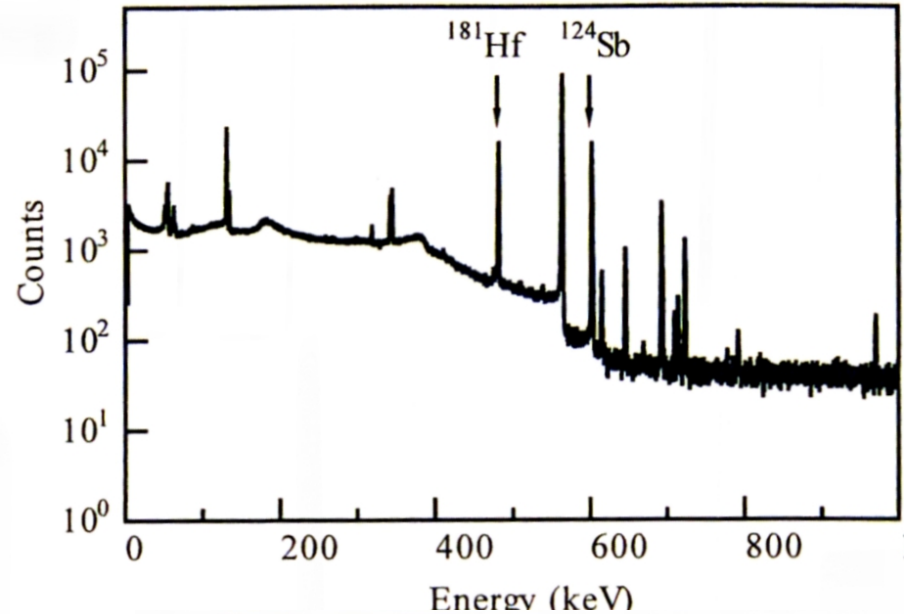
Neutrons for « high-k » dielectrics

Microelectronics standards:
process control at **4 %**

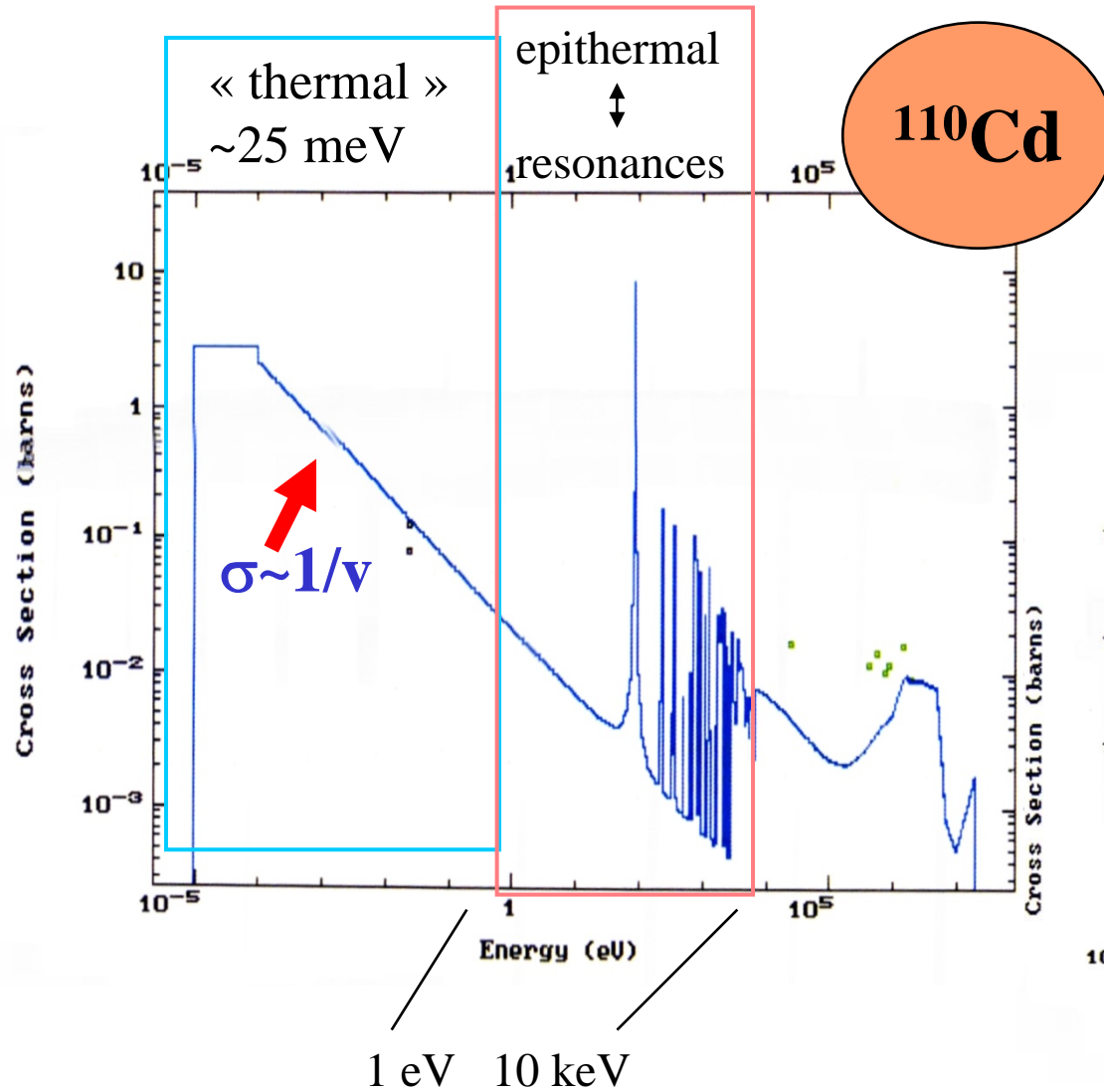


Nice candidate: ^{181}Hf !

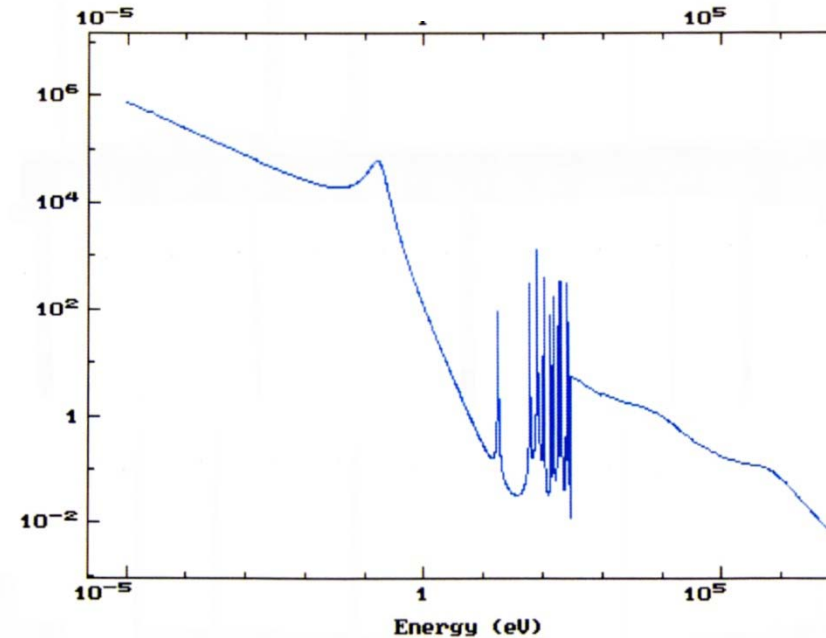
- $\epsilon = 24$ (compare SiO_2 : 3.9)
 - HfO_2 compatible with Si lattice
 - $^{181}\text{Hf} = \gamma$ -emitter (482 keV)
 - Precision peak integration with $^{124}\text{Sb}/\text{Hf}$ ratio (4 h *n* irradi)
- ➔ Obtained precision = 0.92%

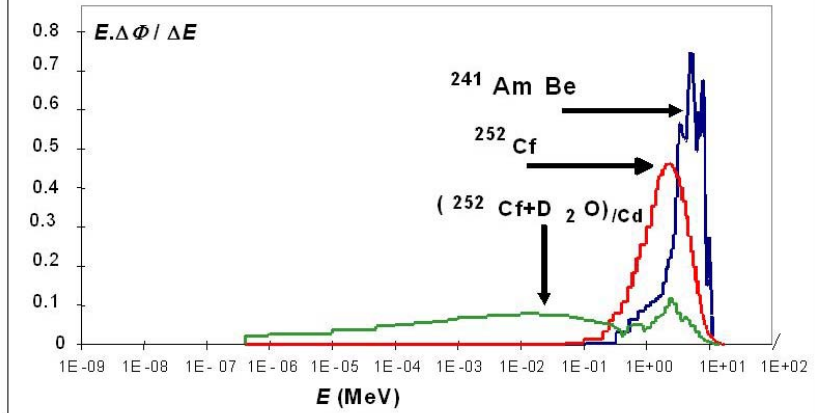


Cadmium, the multipurpose *absorber*



^{113}Cd (12% of natural Cd):
higher σ (thermal region)
but no longer $1/v$ law...

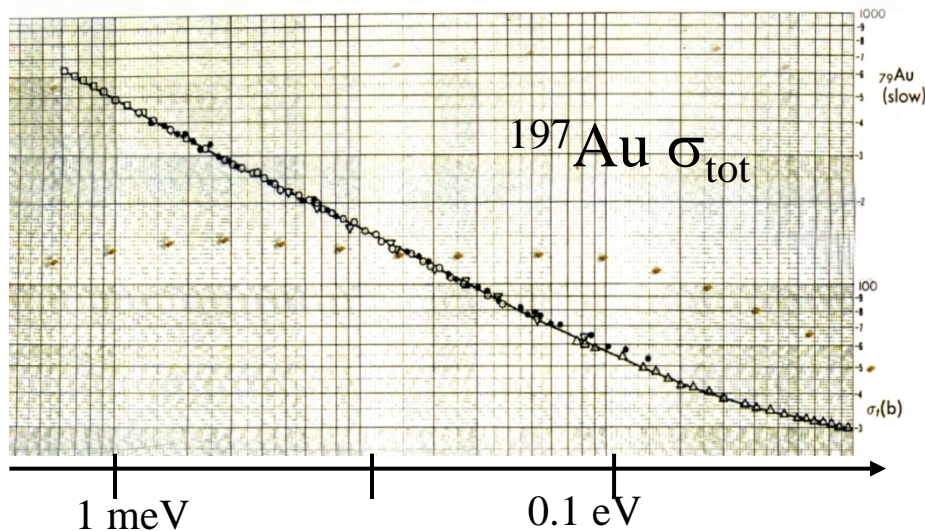




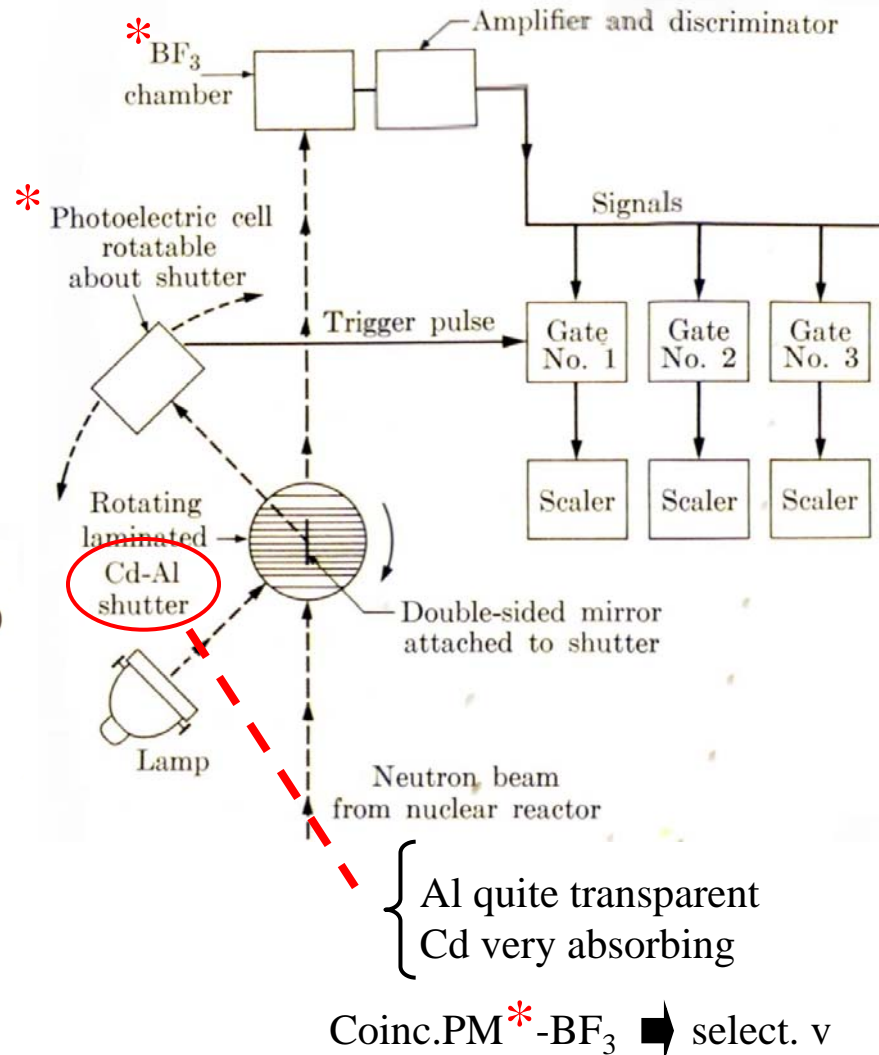
1) modified spectra (slow- n enriched)

2) velocity selectors (« choppers »)

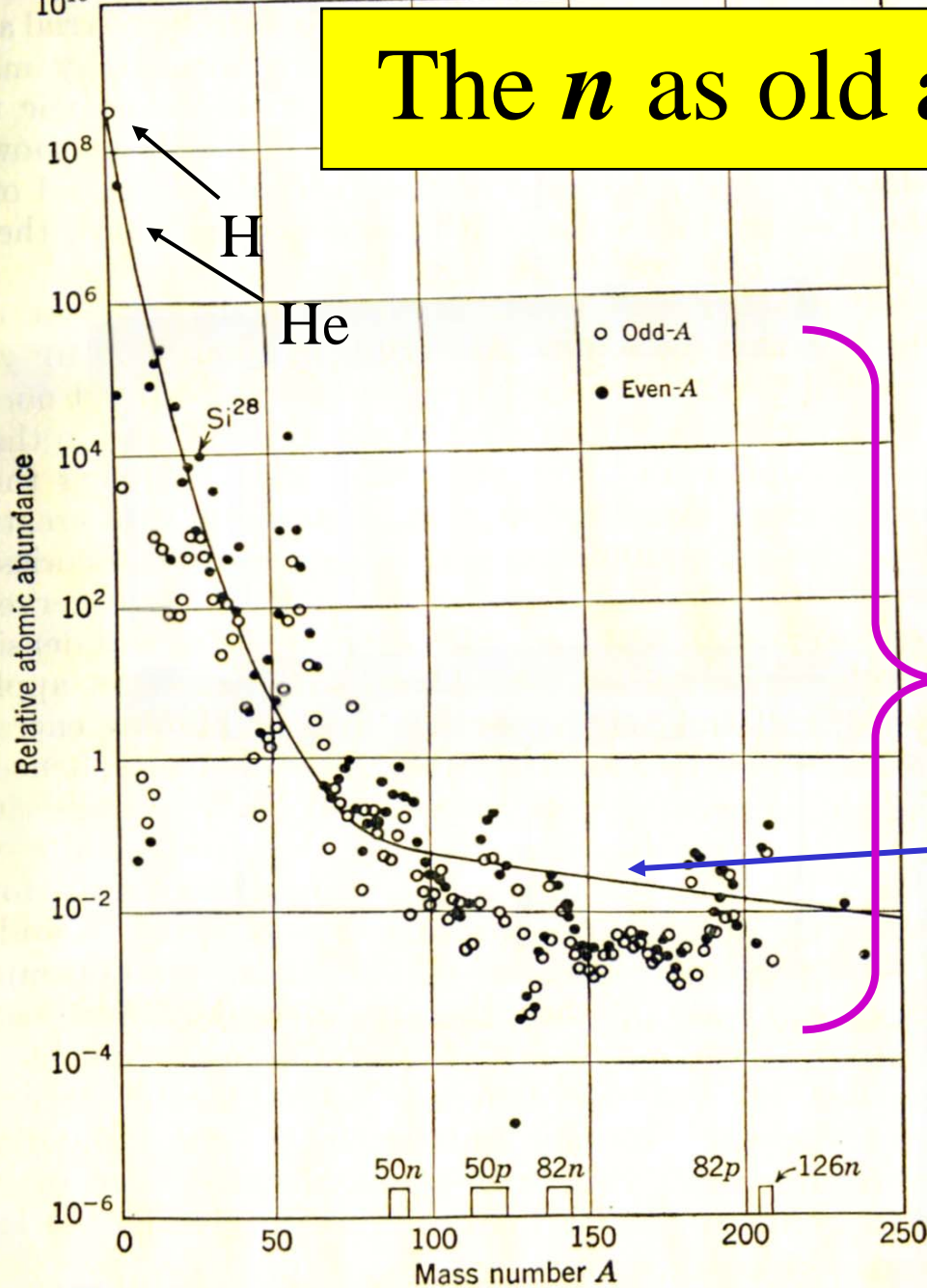
→ nice $\sigma=f(E)$



Applications of Cd



The n as old as the Universe...

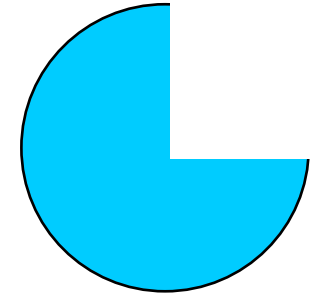


Our universe:

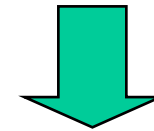
= 75% H

+25% He...

+ small ϵ



This $\epsilon(A)$ can be computed
with pretty nice precision !
(full line = calculation)



What kind of
« calculation »???

« Hadronic era »

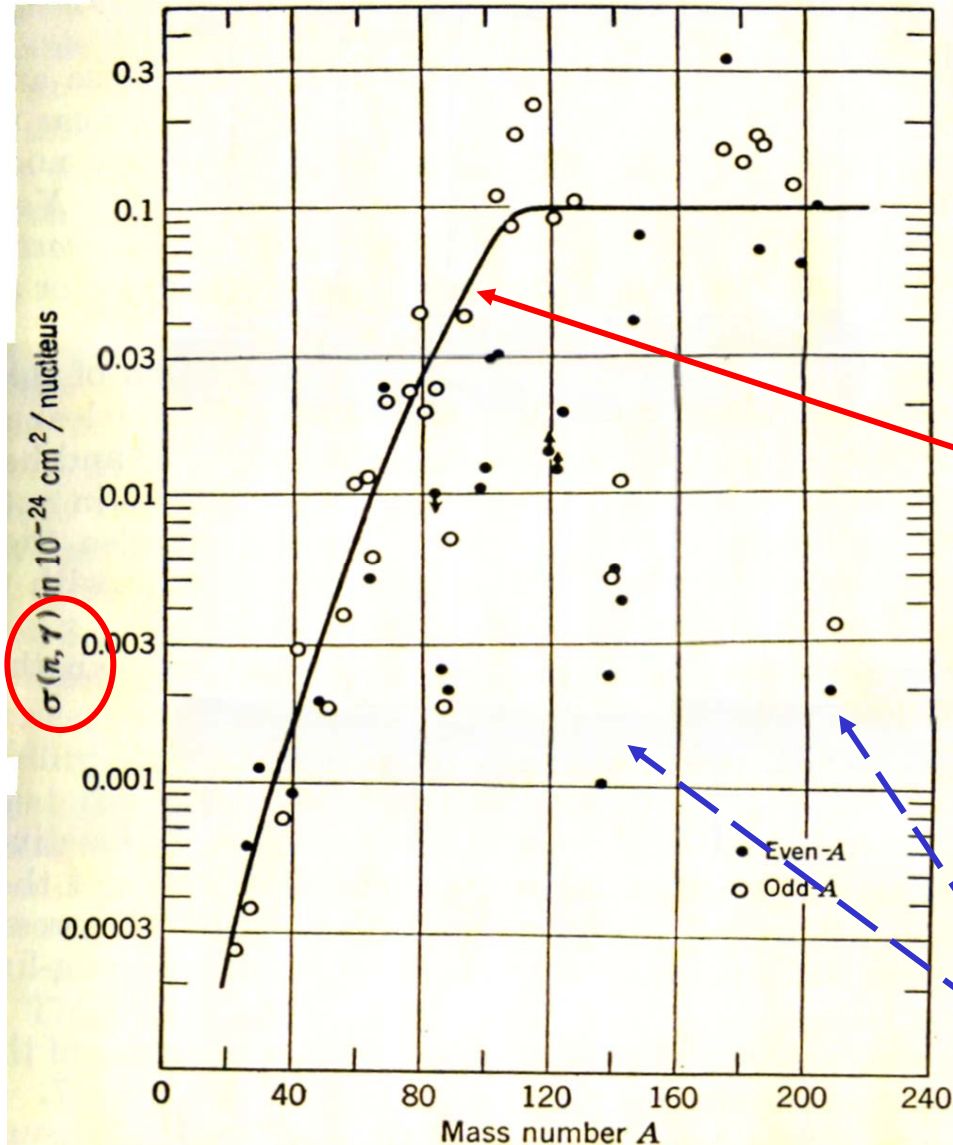
$T^\circ(\text{Univ}) \sim 10^{12}\text{K} \sim \text{MeV}:$

Take a baryonic (p,n)- « soup »
+ « equilibrium » state with π^{+-0}
+ adiabatic expansion ($\Delta t \sim 0.1 \mu\text{s}$)
+ *neutron capture* cross section

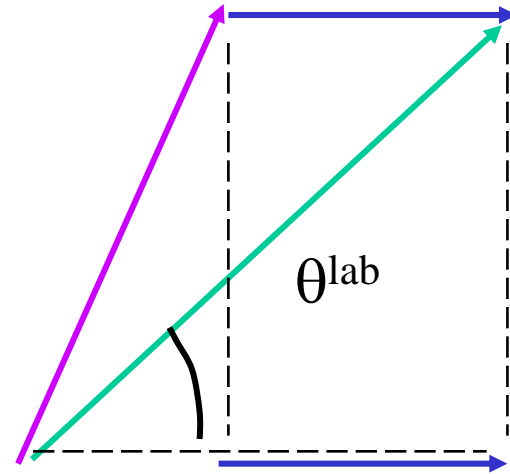
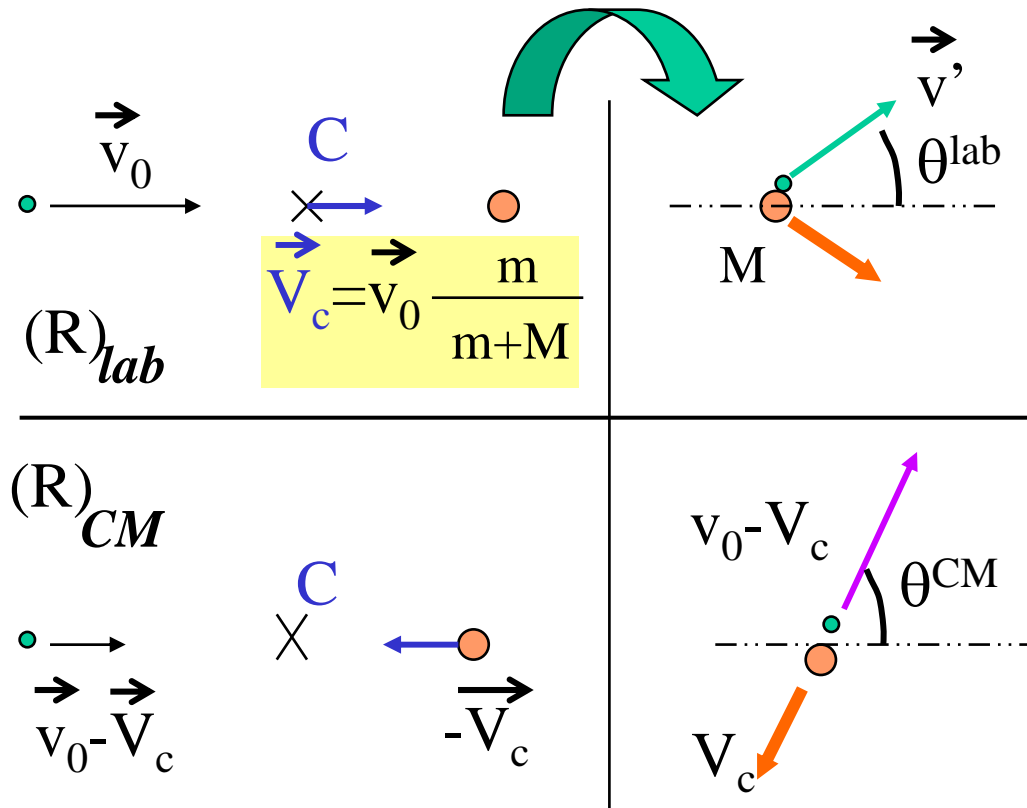
= the *observed* relative abundance of elements !

(=previous plot)

(detail: closed shells => additional neutron)



Changing the reference frame



$$\text{tg } \theta^{lab} = \frac{\sin \theta^{CM} v_0 (M/M+m)}{V_c + \cos \theta^{CM} v_0 (M/M+m)}$$



$$\text{tg } \theta^{lab} = \frac{\sin \theta^{CM}}{m/M + \cos \theta^{CM}}$$

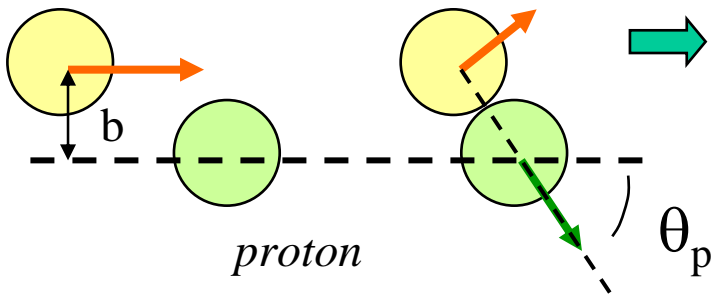
The rules of billiard game (detail)

Elastic (n,p) behaves like *hard spheres* (square nuc.potential !)

$$\sin \theta_p^{\text{LAB}} = b/2R$$

Distributions: *n*-beam uniform \Rightarrow $f(b^2)db^2$ uniform

MC guys!

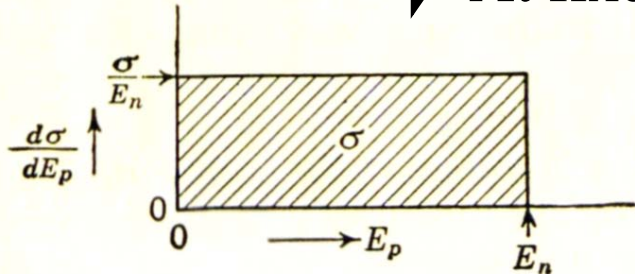


$\sin^2 \theta_p$ uniform (not θ_p , and not $\sin \theta_p$!)

\Rightarrow $\cos^2 \theta_p$ uniform also

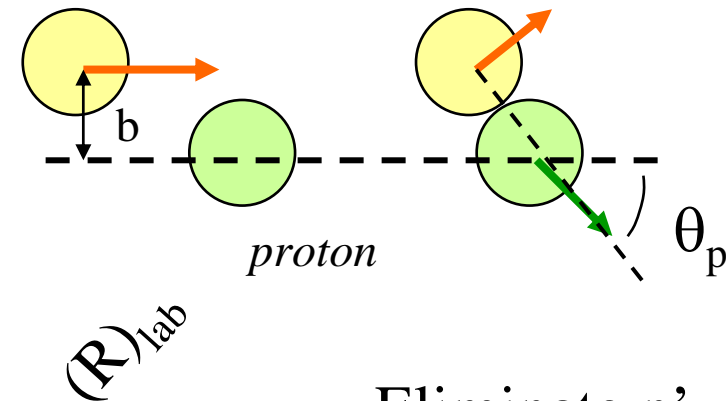
\Rightarrow At fixed E_n : $\left\{ \begin{array}{l} E_p = E_n \cdot \cos^2 \theta_p \text{ distribution is flat !} \\ \text{and } \langle E_p \rangle = E_n / 2 \text{ (LAB)} \end{array} \right.$

(all this is \Leftrightarrow to isotropy in CM...)



The rules of billiard (detail of detail)

Demonstration of $E_p = E_n \cdot \cos^2 \theta_p$



From \vec{p} and $E_n = E_n' + E_p$: $\theta_n + \theta_p = \pi/2$ *

$$\vec{p}_n = \vec{p}_n' + \vec{p}_p \quad : \quad \begin{cases} p_n = p_n' \cos \theta_n + p_p \cos \theta_p \\ 0 = p_n' \sin \theta_n - p_p \sin \theta_p \end{cases}$$

Eliminate n' \Rightarrow
$$p = p_p \left\{ \frac{\sin \theta_p \cos \theta}{\sin \theta} + \cos \theta \right\}$$

$$= p_p \cdot \frac{\sin (\theta_p + \theta)^*}{\sin (\theta)}$$

$$= p_p / \sin (\theta)$$

\Rightarrow
$$p_p = p_n \cdot \sin \theta_n = p_n \cdot \cos \theta_p$$

or
$$\mathbf{E}_p = \mathbf{E}_n \cdot \cos^2 \theta_p \quad (\text{as } m_p = m_n)$$

Behind the « Cloud chamber »

